Document: Math 207C (Spring 2012)
Professor: Hunter
Latest Update: June 6, 2012
Author: Jeff Irion
http://www.math.ucdavis.edu/~jlirion

## Contents

1 4-2-12 ..... 4
1.1 Dimensional Analysis ..... 4
1.2 Fluids Flows, Reynold's Number ..... 4
2 4-4-12 ..... 6
2.1 Navier-Stokes Equation ..... 6
2.2 Low Reynolds Number Flows $(\operatorname{Re} \rightarrow 0)$ ..... 7
2.3 High Reynolds Number Limit $(\operatorname{Re} \rightarrow \infty)$ ..... 7
2.4 Similarity Solutions ..... 7
3 4-6-12 ..... 9
3.1 Heat Equation ..... 9
$4 \quad 4-9-12$ ..... 12
4.1 Heat Equation ..... 12
4.1.1 Fourier Transform ..... 12
4.2 Back to the Heat Equation ..... 13
4.3 A Porous Medium Problem ..... 14
5 4-11-12 ..... 16
5.1 Porous Medium Equation ..... 16
5.2 Perturbation Theory ..... 17
6 4-13-12 ..... 18
6.1 Regular vs. Singular Perturbations ..... 18
6.1.1 Example \#2 ..... 19
7 4-16-12 ..... 21
7.1 Asymptotic and Convergent Series ..... 21
7.1.1 Optimal Truncation ..... 22
7.2 Notation for Asymptotic Behavior ..... 23
8 4-18-12 ..... 24
8.1 Perturbation Theory for ODE's ..... 24
8.2 Overdamped Simple Harmonic Oscillator (Logan 2.4) ..... 24
9 4-20-12 ..... 27
9.1 Strongly Damped Oscillator ..... 27
9.2 Phase Plane ..... 29
9.3 Michaelis Menton Enzyme Kinetics ..... 29
10 4-23-12 ..... 31
10.1 Enzyme Kinetics (Continued) ..... 31
11 4-25-12 ..... 34
11.1 Geometric Singular Perturbation Theory ..... 34
11.2 Van der Pol Oscillator ..... 35
12 4-27-12 ..... 36
12.1 Heat Flow in a Slowly-Varying Rod ..... 36
13 4-30-12 ..... 40
13.1 Boundary Layer Problems ..... 40
13.2 Model Boundary Layer Problem ..... 40
14 5-2-12 ..... 43
14.1 Follow-Up: Why is the boundary layer at 0 ? ..... 43
14.2 General Linear 2nd Order BVP's ..... 43
14.2.1 Boundary Layer Example 1 ..... 44
15 5-4-12 ..... 46
15.1 Boundary Layers (Continued) ..... 46
15.1.1 Boundary Layer Example 1 (From Last Time) ..... 46
15.1.2 Boundary Layer Example 2 ..... 47
16 5-7-12 ..... 49
16.1 Boundary Layer Example 2 ..... 49
16.2 Boundary Layer Example 3 ..... 51
17 5-9-12 ..... 52
17.1 Boundary Layer Example 3 ..... 52
18 5-11-12 ..... 55
18.1 Method of Multiple Scales (MMS) and Oscillations ..... 55
18.1.1 Example ..... 56
18.2 Poincaré-Lindstedt Method ..... 56
19 5-14-12 ..... 57
19.1 Poincaré-Lindstedt Method ..... 57
20 5-16-12 ..... 60
20.1 Poincaré-Lindstedt Method ..... 60
20.2 Weakly Damped Simple Harmonic Oscillator ..... 61
21 5-18-12 ..... 63
21.1 Weakly Damped Oscillator ..... 63
21.2 van der Pol Oscillator ..... 64
22 5-21-12 ..... 65
22.1 van der Pol Equation ..... 65
23 5-23-12 ..... 67
23.1 Method of Averaging ..... 67
23.2 Geometrical Interpretation ..... 68
23.3 Periodic Standard Form ..... 69
24 5-25-12 ..... 70
24.1 WKB Method ..... 70
25 5-30-12 ..... 72
25.1 WKB Method ..... 72
25.2 Schrödinger Equation ..... 74
26 6-1-12 ..... 75
26.1 WKB Method and Turning Points ..... 75
26.2 A Model Bifurcation Problem for PDEs ..... 76
27-6-4-12 ..... 78
27.1 Model PDE Bifurcation Problem ..... 78
28 6-6-12 ..... 81
28.1 Outline of Topics ..... 81
28.2 Sample Problems ..... 82

### 1.1 Dimensional Analysis

We have a fundamental system of units: $\left(d_{1}, d_{2}, \ldots, d_{r}\right)$.

## Example 1.1. Mechanics

- mass $M$
- length $L$
- time $T$

Derived units, e.g. velocity $V=\frac{L}{T}$ and acceleration $A=\frac{L}{T^{2}}$. We can use different sets of units as fundamental units (provided they're independent). For example, we could use mass, velocity, and acceleration. Any model of a system must be invariant under rescalings that correspond to changes in the system of units.

Let's say we have a fundamental system of (independent) units: $d_{1}, d_{2}, \ldots, d_{r}$. We have a set of quantities in the model:

$$
\left\{\begin{array}{rr}
a_{1}, a_{2}, \ldots, a_{r} & \text { with dimension }\left[a_{i}\right]=d_{i} \\
\vdots & \\
b_{r+1}, \ldots, b_{n} &
\end{array}\right.
$$

Let's say $b_{j}$ has dimensions

$$
\left[b_{j}\right]=d_{1}^{\beta_{1 j}} d_{2}^{\beta_{2 j}} \cdots d_{r}^{\beta_{r j}} .
$$

Then the model can only depend on

$$
\Pi_{j}=\frac{b_{j}}{a_{1}^{\beta_{1 j}} a_{2}^{\beta_{2 j}} \cdots a_{r}^{\beta_{r j}}} .
$$

So our model has:

- $r$ independent dimensions
- $n$ independent quantities

Then dimensional analysis says it depends on $n-r$ dimensionless variables. (This is called the Buckingham Pi Theorem.)

### 1.2 Fluids Flows, Reynold's Number

Let's say we have a sphere in a flow. What is the drag on the sphere?
Parameters:

- $u=$ speed of the fluid, $[u]=\frac{L}{T}$
- $d=$ diameter of the sphere, $[d]=L$
- $\mu=$ viscosity of the fluid, $[\mu]=\frac{M}{L T}$
- $\rho_{0}=$ density of the fluid, $\left[\rho_{0}\right]=\frac{M}{L^{3}}$
- Assume the fluid is incompressible (this is OK if $u \ll c_{0}$, the speed of sound in the fluid)

Fundamental units: $M, L, T$.
In a Newtonian fluid:

- $T=$ viscous stress tensor,

$$
T=\mu\left(\nabla u+\nabla u^{T}\right),
$$

where $u=$ velocity. This gives the force/unit area. The dimensions of $T$ are

$$
\begin{aligned}
{[T] } & =\frac{M L}{T^{2}} \cdot \frac{1}{L^{2}}=\frac{M}{L T^{2}} \\
{[\nabla u] } & =\frac{1}{T} \\
{[\mu] } & =\frac{M}{L T}
\end{aligned}
$$

We define the kinematic viscosity:

$$
\begin{aligned}
\nu & =\frac{\mu}{\rho_{0}} \\
{[\nu] } & =\frac{L^{2}}{T}
\end{aligned}
$$

The physical interpretation of this quantity is diffusivity of momentum.

$$
\begin{aligned}
& \nu \approx 1 \mathrm{~mm}^{2} / \mathrm{s} \text { in water } \\
& \nu \approx 15 \mathrm{~mm}^{2} / \mathrm{s} \text { in air }
\end{aligned}
$$

We can define the Reynold's number:

$$
\operatorname{Re}=\frac{u d}{\nu} .
$$

This is the crucial dimensionless parameter that controls everything.
Back to our question about drag on a sphere. $D=$ drag force with dimensions $[D]=\frac{M L}{T^{2}}$.

$$
\begin{aligned}
{\left[\rho_{0} u^{2} d^{2}\right] } & =\frac{M}{L^{3}} \cdot \frac{L^{2}}{T^{2}} \cdot L^{2}=\frac{M L}{T^{2}} \\
\frac{D}{\rho_{0} u^{2} d^{2}} & =F(\operatorname{Re}) \\
D & =\rho_{0} u^{2} d^{2} F(\mathrm{Re})
\end{aligned}
$$

## 2 4-4-12

### 2.1 Navier-Stokes Equation

$$
\begin{aligned}
\rho_{0}\left(\vec{u}_{t}+\vec{u} \cdot \nabla \vec{u}\right)+\nabla p & =\mu_{0} \Delta \vec{u} \\
\nabla \cdot \vec{u} & =0
\end{aligned}
$$

- $\vec{u}=\vec{u}(\vec{x}, t)$ is the fluid velocity
- $p=p(\vec{x}, t)$ is the pressure
- $\vec{u}=\left(u_{1}, u_{2}, u_{3}\right)$
- $\vec{x}=\left(x_{1}, x_{2}, x_{3}\right)$
- $\nabla=\left(\frac{\partial}{\partial x_{1}}, \frac{\partial}{\partial x_{2}}, \frac{\partial}{\partial x_{3}}\right)$

Parameters

- $\rho_{0}=$ fluid density
- $\mu=$ fluid viscosity
- $U=$ "typical" flow velocity
- $L=$ "typical" flow length scale

Dimensionless variables

- $\vec{u}^{*}=\frac{\vec{u}}{U}$
- $\vec{x}^{*}=\frac{\vec{x}}{L}$
- $t^{*}=\frac{U t}{L}$
- $p^{*}=\frac{p}{\rho_{0} U^{2}}$
$-[\nabla p]=\left[\rho_{0} \vec{u}_{t}\right]$
$-\frac{[p]}{L}=\left[\rho_{0}\right] \frac{L}{T^{2}}$
$-[p]=\left[\rho_{0}\right] \frac{L^{2}}{T^{2}}$
- $\nabla=\frac{1}{L} \nabla^{*}$
- $\partial_{t}=\frac{d t^{*}}{d t} \partial_{t^{*}}=\frac{U}{L} \partial_{t^{*}}$

$$
\begin{aligned}
\rho_{0}\left[\frac{U}{L}\left(U \vec{u}^{*}\right)_{t^{*}}+\frac{U^{2}}{L} \vec{u}^{*} \cdot \nabla^{*} \vec{u}^{*}\right]+\frac{\rho_{0} U^{2}}{L} \nabla^{*} p^{*} & =\frac{\mu U}{L^{2}} \Delta^{*} \vec{u}^{*} \\
\vec{u}_{t^{*}}^{*}+\vec{u}^{*} \cdot \nabla^{*} \vec{u}^{*}+\nabla^{*} p^{*} & =\frac{1}{\operatorname{Re}} \Delta^{*} \vec{u}^{*} \\
\nabla^{*} \cdot \vec{u}^{*} & =0
\end{aligned}
$$

### 2.2 Low Reynolds Number Flows ( $\operatorname{Re} \rightarrow 0$ )

$$
\begin{aligned}
p^{*} & =\frac{\tilde{p}}{\operatorname{Re}} \\
\tilde{p} & =\operatorname{Re} \cdot p^{*}=\frac{U L}{\nu} \cdot \frac{p}{\rho_{0} U^{2}}=\frac{L}{\mu U} p
\end{aligned}
$$

As $\operatorname{Re} \rightarrow 0$, we get Stokes equations:

$$
\begin{aligned}
\nabla^{*} \tilde{p} & =\Delta^{*} \vec{u}^{*} \\
\nabla^{*} \cdot \vec{u}^{*} & =0 .
\end{aligned}
$$

These are linear!
Example 2.1. Drag on a Sphere as Re $\rightarrow 0$

$$
D=\rho_{0} U^{2} L^{2} F(\mathrm{Re})
$$

Consider $\lim _{\mathrm{Re} \rightarrow 0} D$. Since the drag is linear in $U$, we need

$$
\begin{aligned}
F(\operatorname{Re}) & =\frac{c}{\operatorname{Re}} \\
D & =\rho_{0}^{2} U^{2} L^{2} \cdot \frac{c}{\operatorname{Re}}=c \frac{\rho_{0} U^{2} L^{2} \nu}{U L}=c \mu_{0} U L
\end{aligned}
$$

Stokes (1851):

$$
D=6 \pi \mu_{0} a U,
$$

where $a$ is the radius of a sphere.

### 2.3 High Reynolds Number Limit $(\operatorname{Re} \rightarrow \infty)$

Formally, we get the Euler equations.

$$
\begin{aligned}
\vec{u}_{t^{*}}^{*}+\vec{u}^{*} \cdot \nabla \vec{u}^{*}+\nabla^{*} p^{*} & =0 \\
\nabla^{*} \cdot \vec{u}^{*} & =0
\end{aligned}
$$

This is nonlinear!
Turbulence, Prandtt boundary layer term $\rightarrow$ singular perturbation neglecting higher derivatives

### 2.4 Similarity Solutions

Consider the heat flow due to a point source.

$$
\begin{aligned}
u_{t} & =v \Delta u \\
u(x, 0) & =E \delta(x)
\end{aligned}
$$

$u(x, t)=$ temperature of (infinite) body. Inject total heat energy $E$ at $x=0$ at $t=0$.

- $\theta=$ temperature dimension, $[u]=\theta$
- $L=$ length, $[x]=L$
- $T=$ time, $[t]=T$

Parameters $\nu, E$

- $[\nu]=\frac{L^{2}}{T}$
- $[E]=\theta L^{n}$
- At $t=0, \int u d x=\int E \delta(x) d x=E$
$-[E]=\left[\int u d x\right]=\theta L^{n}$


### 3.1 Heat Equation

$$
\begin{aligned}
u_{t} & =\nu \Delta u \\
u(x, 0) & =E \delta(x)
\end{aligned}
$$

$u(x, t)$ is the temperature, $x \in \mathbb{R}^{n}$.

## Parameters

- $\nu$ : thermal diffusivity, $[\nu]=\frac{L^{2}}{T}$
- $E$ : initial heat, $[E]=\theta L^{n}$

Dependent variables: $u([u]=\theta)$.
Independent variables: $r([r]=L), t([t]=T)$.
So we have

- 5 quantities: $\nu, E, u, r, t$
- 3 dimensions: $\theta, L, T$

We can form 2 dimensionless quantities.

- Time: $t$
- There is 1 variable with dimensions of time: $t$. This will lead to the self-similarity of the problem. That is, a solution on one time scale is a rescale of a solution on another time scale.
- Length: $\sqrt{\nu t}$
- Temperature: $\frac{E}{\sqrt{\nu t}}$

So we have

$$
\begin{aligned}
u^{*} & =\frac{u}{E /(\nu t)^{n / 2}} \\
u & =\frac{E}{(\nu t)^{n / 2}} u^{*}(\xi) \\
\xi & =\frac{r}{\sqrt{\nu t}}
\end{aligned}
$$

So our dimensionless temperature depends only on $\xi=\frac{r}{\sqrt{\nu t}}$.

Let $u^{*}=F$. We will look for solutions of the form

$$
\begin{aligned}
u & =\frac{E}{(\nu t)^{n / 2}} F\left(\frac{r}{\sqrt{\nu t}}\right) \\
u_{t} & =\nu \frac{1}{r^{n-1}} \frac{\partial}{\partial r}\left(r^{n-1} \frac{\partial u}{\partial r}\right) \\
u_{t} & =\frac{\left(-\frac{n}{2}\right) E}{\nu^{n / 2} t^{\frac{n}{2}+1}} F+\frac{E}{(\nu t)^{n / 2}} F^{\prime}\left(\frac{r}{\sqrt{\nu t}}\right)\left(-\frac{1}{2}\right) \frac{r}{\sqrt{\nu} t^{3 / 2}} \\
u_{t} & =\frac{-E}{\nu^{n / 2} t^{\frac{n}{2}+1}}\left[\frac{n}{2} F+\frac{1}{2} F^{\prime} \frac{r}{\sqrt{\nu t}}\right] \\
& =-\frac{E}{\nu^{n / 2} t^{\frac{n}{2}+1}}\left[\xi F^{\prime}+n F\right] \\
\Delta u & =\frac{E}{(\nu t)^{n / 2}} \frac{1}{r^{n-1}} \frac{\partial}{\partial r}\left(r^{n-1} \frac{\partial F}{\partial r}\right) \\
& =\frac{E}{(\nu t)^{\frac{n}{2}+1}} \frac{1}{\xi^{n-1}} \frac{d}{d \xi}\left(\xi^{n-1} \frac{d F}{d \xi}\right) \\
-\frac{1}{2} \frac{E}{\chi^{n / 2} t^{\frac{n}{2}+1}}\left[\xi F^{\prime}+n F\right] & =\nu \frac{E}{(\nu t)^{\frac{n}{2}+1}} \frac{1}{\xi^{n-1}} \frac{d}{d \xi}\left(\xi^{n-1} \frac{d F}{d \xi}\right) \\
\frac{1}{\xi^{n-1}} \frac{d}{d \xi}\left(\xi^{n-1} \frac{d F}{d \xi}\right) & =-\frac{1}{2}\left(\xi F^{\prime}+n F\right)
\end{aligned}
$$

So we have reduced our PDE to an ODE for $F(\xi)$. This is a second-order, variable coefficient ODE. We have

$$
\begin{aligned}
F^{\prime \prime}+\frac{n-1}{\xi} F^{\prime} & =-\frac{1}{2} \xi F^{\prime}-\frac{1}{2} n F \\
F^{\prime \prime}+\left(\frac{n-1}{\xi}+\frac{1}{2} \xi\right) F^{\prime}+\frac{1}{2} n F & =0 \\
\underbrace{\left(F^{\prime}+\frac{1}{2} \xi F\right)}_{G}+\frac{n-1}{\xi}\left(F^{\prime}+\frac{1}{2} \xi F\right) & =0 \\
\xi^{n-1} G^{\prime}+(n-1) \xi^{n-2} G & =0 \\
\left(\xi^{n-1} G\right)^{\prime} & =0 \\
G & =\frac{c}{\xi^{n-1}}
\end{aligned}
$$

Take $c=0$; otherwise $G \rightarrow \infty$ as $\xi \rightarrow 0(r \rightarrow 0)$. So

$$
\begin{aligned}
G & =0 \\
F^{\prime}+\frac{1}{2} \xi F & =0 \\
\left(e^{\xi^{2} / 4} F\right)^{\prime} & =0 \\
e^{\xi^{2} / 4} F & =c \quad(\text { constant }) \\
F(\xi) & =c e^{-\xi^{2} / 4}
\end{aligned}
$$

Using the initial condition:

$$
\begin{aligned}
\int u(x, 0) d x & =E \\
\Rightarrow \quad c & =\frac{1}{(4 \pi)^{n / 2}} \\
u(x, t) & =\frac{E}{(4 \pi \nu t)^{n / 2}} \exp \left(-\frac{|x|^{2}}{4 \nu t}\right)
\end{aligned}
$$



### 4.1 Heat Equation

$$
\begin{aligned}
u_{t} & =\nu \Delta u \\
u(x, 0) & =E \delta(x)
\end{aligned}
$$

Since this is a linear PDE with constant coefficients (on $\mathbb{R}^{n}$ ), we can solve this using the Fourier transform.

### 4.1.1 Fourier Transform

$$
\begin{aligned}
& f(x), \quad x=\left(x_{1}, \ldots, x_{n}\right) \in \mathbb{R}^{n} \\
& \hat{f}(k), \quad k=\left(k_{1}, \ldots, k_{n}\right) \in \mathbb{R}^{n} \\
& \hat{f}(k)=\frac{1}{(2 \pi)^{n}} \int_{\mathbb{R}^{n}} f(x) e^{-i k \cdot x} d x \\
& f(x)=\int_{\mathbb{R}^{n}} \hat{f}(k) e^{i k \cdot x} d k
\end{aligned}
$$

We say that $\hat{f}=\mathcal{F}[f]$, where $\mathcal{F}$ is the Fourier transform. Then

$$
\begin{aligned}
\frac{\partial f}{\partial x_{\alpha}}(x) & =\frac{\partial}{\partial x_{\alpha}} \int \hat{f}(k) e^{i k \cdot x} d k \\
& =\int \hat{f}(k) \frac{\partial}{\partial x_{\alpha}}\left(e^{i k \cdot x}\right) d k \\
& =\int i k_{\alpha} \hat{f}(k) e^{i k \cdot x} d k \\
\mathcal{F}\left(\frac{\partial f}{\partial x_{\alpha}}\right) & =i k_{\alpha} \hat{f}(k)
\end{aligned}
$$

In particular,

$$
\mathcal{F}[\Delta f]=-|k|^{2} \hat{f}(k)
$$

We can define $\sqrt{-\Delta}$ by

$$
\mathcal{F}[\sqrt{-\Delta} f]=|k| \hat{f}(k)
$$

Example 4.1.

$$
\begin{aligned}
& f(x)=e^{-|x|^{2} / 2 \sigma^{2}} \\
& \hat{f}(k)=\left(\frac{\sigma}{\sqrt{2 \pi}}\right)^{n} e^{-\sigma^{2}|k|^{2} / 2}
\end{aligned}
$$



### 4.2 Back to the Heat Equation

$$
\begin{aligned}
u(x, t) & =\int_{\mathbb{R}^{n}} \hat{u}(k, t) e^{i k \cdot x} d k \\
\hat{u} & =\mathcal{F}[u] \\
\mathcal{F}\left[u_{t}\right] & =\hat{u}_{t} \\
\mathcal{F}[\Delta u] & =-|k|^{2} \hat{u} \\
\mathcal{F}[\delta(x)] & =\frac{1}{(2 \pi)^{n}} \int \delta(x) e^{-i k \cdot x} d x=\frac{1}{(2 \pi)^{n}}
\end{aligned}
$$

So the heat equation becomes

$$
\begin{aligned}
\hat{u}_{t} & =-\nu|k|^{2} \hat{u} \\
\hat{u}(k, 0) & =\frac{E}{(2 \pi)^{n}}
\end{aligned}
$$

The solutions look like

$$
\hat{u}(k, t)=\frac{E}{(2 \pi)^{n}} e^{-\nu|k|^{2} t}
$$



$$
u(x, t)=\frac{E}{(4 \pi \nu t)^{n / 2}} e^{-|x|^{2} / 4 \nu t}
$$



Figure 1: The heat diffuses with time.

This is a Green's function:

$$
\begin{gathered}
G(x, t)=\frac{1}{(4 \pi \nu t)^{n / 2}} e^{-|x|^{2} / 4 \nu t} \\
G_{t}=\nu \Delta G \\
G(x, 0)=\delta(x)
\end{gathered}
$$

So the solution of the heat equation,

$$
\begin{aligned}
u_{t} & =\nu \Delta u \\
u(x, 0) & =f(x)
\end{aligned}
$$

is

$$
\begin{aligned}
u(x, t) & =\int_{\mathbb{R}^{n}} G(x-\xi, t) f(\xi) d \xi \\
& =\frac{1}{(4 \pi \nu t)^{n / 2}} \int_{\mathbb{R}^{n}} \exp \left(-\frac{|x-\xi|^{2}}{4 \nu t}\right) f(\xi) d \xi
\end{aligned}
$$

### 4.3 A Porous Medium Problem



Figure 2: The aquifer is fully saturated with water. $z=h(x, t)$ is the height of the aquifer.
Assume slow transverse flow, so the pressure is hydrostatic:

$$
p=\rho g(h-z)
$$

The pressure head is

$$
\begin{aligned}
& H=p+\rho g z \\
& H=\rho g h \quad \text { independent of } z .
\end{aligned}
$$

Assume the fluid is incompressible $\Rightarrow$ conservation of volume. The change in the volume between $a$ and $b$ is

$$
\begin{align*}
\frac{d}{d t} \int_{a}^{b} h d x & =-[h v]_{x=a}^{x=b} \\
& =-\int_{a}^{b}(h v)_{x} d x \\
\int_{a}^{b}\left[h_{t}+(h v)_{x}\right] d x & =0 \quad \forall[a, b] \\
h_{t}+(h v)_{x} & =0 . \tag{4.1}
\end{align*}
$$

Darcy's law:

$$
v=-\frac{k}{\mu} \nabla H
$$

$k$ is the permeability, and $\mu$ is the fluid viscosity. This is saying that the velocity is proportional to the gradient of the pressure head. In our case, we have

$$
v=-\frac{k}{\mu} \rho g h_{x} .
$$

Plugging this into (4.1), we get

$$
\begin{aligned}
h_{t} & =K\left(h h_{x}\right)_{x} \\
K & =\frac{k \rho g}{\mu}
\end{aligned}
$$

This is the 1D porous medium equation. This is a nonlinear, degenerate diffusion equation. When $h \rightarrow 0$, the diffusion drops out.

## $5 \quad 4-11-12$

### 5.1 Porous Medium Equation

$$
\begin{aligned}
h_{t} & =k\left(h h_{x}\right)_{x} \\
h(x, 0) & =I \delta(x)
\end{aligned}
$$

(Barenblatt)
Dimensions

- (vertical) height $H$
- (horizontal) length $L$
- time $T$

Dependent Variables: $h(H)$
Independent Variables: $x(L), t(T)$
Parameters: $k\left(\frac{L^{2}}{H T}\right), I(H L)$
Use $t, k, I$ to nondimensionalize the problem.

$$
\begin{aligned}
{[t] } & =T \\
{\left[(k I t)^{1 / 3}\right] } & =L \\
{\left[\frac{I}{(k I t)^{1 / 3}}\right] } & =H \\
h(x, t) & =\frac{I^{2 / 3}}{(k t)^{1 / 3}} F\left(\frac{x}{(k I t)^{1 / 3}}\right) \\
\int h(x, t) d x & =I \int F(\xi) d \xi \\
-\frac{1}{3} \frac{I^{2 / 3}}{k^{1 / 3} t^{4 / 3}} F+\frac{I^{2 / 3}}{(k t)^{1 / 3}}\left(-\frac{1}{3}\right) \frac{x}{(k I)^{1 / 3} t^{4 / 3}} F^{\prime} & \\
& =k\left[\frac{I^{2 / 3}}{(k t)^{1 / 3}}\right]^{2} \frac{1}{(k I t)^{2 / 3}}\left(F F^{\prime}\right)^{\prime} \\
-\frac{1}{3} F-\frac{1}{3} \xi F^{\prime} & =\left(F F^{\prime}\right)^{\prime}, \\
\left(F F^{\prime}\right)^{\prime} & =-\frac{1}{3}\left(\xi F^{\prime}+F\right) \\
& =-\frac{1}{3}(\xi F)^{\prime} \\
F F^{\prime} & =-\frac{1}{3} \xi F+c
\end{aligned}
$$

We expect $F \rightarrow 0$ as $\xi \rightarrow \infty$. Take $c=0$.

$$
\begin{aligned}
F F^{\prime} & =-\frac{1}{3} \xi F \\
F^{\prime} & =-\frac{1}{3} \xi \\
F(\xi) & =\frac{1}{6}\left(a^{2}-\xi^{2}\right)
\end{aligned}
$$

We need

$$
\begin{aligned}
& \int_{-\infty}^{\infty} F(\xi) d \xi=1 \\
& F(\xi)=\left\{\begin{array}{rr}
\frac{1}{6}\left(a^{2}-\xi^{2}\right) & |\xi|<a \\
0 & |\xi| \geq a
\end{array}\right. \\
& \int_{-a}^{a} \frac{1}{6}\left(a^{2}-\xi^{2}\right) d \xi=1 \\
& a=\left(\frac{9}{2}\right)^{1 / 3} \\
& h(x, t)= \begin{cases}\frac{I^{2 / 3}}{6(k t)^{1 / 3}}\left[\left(\frac{9}{2}\right)^{2 / 3}-\frac{x^{2}}{(k I t)^{2 / 3}}\right] \\
0 & |x|<\left(\frac{9 k I t}{2}\right)^{1 / 3} \\
\text { otherwise }\end{cases} \\
& \underbrace{}_{t^{1 / 3}}
\end{aligned}
$$

### 5.2 Perturbation Theory

$$
p^{\epsilon}(x)=0
$$

Problem for $x$ depending on a small parameter $\epsilon$. Solution:

$$
x=x(\epsilon)
$$

Suppose $p^{\epsilon}$ "simplifies" at $\epsilon=0$. Goal: to find approximations of the solution $x(\epsilon)$ when $\epsilon$ is small.

## Definition 5.1. Regular, Singular

Classify perturbation problem as

- regular if the $\epsilon=0$ problem is "close" to the $\epsilon \neq 0$ problem
- singular if the $\epsilon=0$ problem is "different" from the $\epsilon \neq 0$ problem


## $6 \quad 4-13-12$

### 6.1 Regular vs. Singular Perturbations

Example 6.1.

$$
x^{3}-x+\epsilon=0
$$

Look for a solutions

$$
\begin{aligned}
& x(\epsilon)=x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots \\
& x^{3}=\left(x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots\right)^{3} \\
& =x_{0}^{3}+3 \epsilon x_{0}^{2} x_{1}+\epsilon^{2}\left[3 x_{0}^{2} x_{2}+3 x_{0} x_{1}^{2}\right]+\cdots \\
& x_{0}^{3}+3 \epsilon x_{0}^{2} x_{1}+\epsilon^{2}\left[3 x_{0}^{2} x_{2}+3 x_{0} x_{1}^{2}\right]+\cdots-x_{0}-\epsilon x_{1}-\epsilon^{2} x_{2}-\cdots+\epsilon=0 \\
& x_{0}^{3}-x_{0}=0 \\
& 3 x_{0}^{2} x_{1}-x_{1}+1=0 \\
& 3 x_{0}^{2} x_{2}+-x_{2}+3 x_{0} x_{1}^{2}=0 \\
& x_{0}=0, \pm 1 \\
& x_{1}=\frac{1}{1-3 x_{0}^{2}} \\
& x_{2}=\frac{3 x_{0} x_{1}^{2}}{1-3 x_{0}^{2}} \\
& x_{0}=0: \quad x=0+\epsilon+0 \cdot \epsilon^{2}+O\left(\epsilon^{3}\right) \\
& x_{0}=1: \quad x=1-\frac{1}{2} \epsilon-\frac{3}{8} \epsilon^{2}+O\left(\epsilon^{3}\right) \\
& x_{0}=-1: \quad x=-1-\frac{1}{2} \epsilon+\frac{3}{8} \epsilon^{2}+O\left(\epsilon^{3}\right)
\end{aligned}
$$

### 6.1.1 Example \#2

$$
\begin{aligned}
\epsilon x^{3}-x+1 & =0 \\
x & =x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots \\
\epsilon\left(x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots\right)^{3}-\left(x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots\right)+1 & =0 \\
\epsilon x_{0}^{3}+3 \epsilon^{2} x_{0}^{2} x_{1}+\cdots-x_{0}-\epsilon x_{1}-\epsilon^{2} x_{2}+1 & =O\left(\epsilon^{3}\right) \\
-x_{0}+1 & =0 \\
x_{0}^{3}-x_{1} & =0 \\
3 x_{0}^{2} x_{1}-x_{2} & =0 \\
x_{0} & =1 \\
x_{1} & =1 \\
x_{2} & =3 \\
x & =1+\epsilon+3 \epsilon^{2}+\cdots
\end{aligned}
$$

This equation is singular: the cubic equation degenerates to a linear equation at $\epsilon=0$.
We only get one root; the other two go off to $\infty$ as $\epsilon \rightarrow 0$. So we introduce a scaled variable:

$$
\begin{aligned}
x & =\frac{y}{\delta(\epsilon)}, \quad y=O(1) \\
\underbrace{\frac{\epsilon}{\delta^{3}} y^{3}}_{(1)}-\underbrace{\frac{1}{\delta} y}_{(2)}+\underbrace{1}_{(3} & =0
\end{aligned}
$$

To get a nontrivial limit, we need a dominant balance between (at least) two terms.

## Two-Term Balances

- (1) $\sim(2): ~ \epsilon / \delta^{3}=1 / \delta ; \delta=\epsilon^{1 / 2}$; (3) $\sim 1$; (1), (2) $\sim 1 / \epsilon^{1 / 2}$; (1) $\sim(2) \gg(3)$
- (2) $\sim(3): 1 / \delta=1 ; \delta=1 ;(2),(3) \sim 1 ;, 1 \gg(1) \sim \epsilon$
- (3) $\sim(1): ~ \epsilon / \delta^{3}=1 ; \delta=\epsilon^{1 / 3}$; (3), (1) $\sim 1 ; 1 \ll(2) \sim 1 / \epsilon^{1 / 3}$

The first two are dominant balances.
To get the remaining roots... $\delta=\epsilon^{1 / 2}$

$$
\begin{aligned}
x & =\frac{y}{\epsilon^{1 / 2}} \\
\frac{\epsilon}{\epsilon^{3 / 2}} y^{3}-\frac{1}{\epsilon^{1 / 2}} y+1 & =0 \\
y^{3}-y+\epsilon^{1 / 2} & =0 \\
y & =y_{0}+\epsilon^{1 / 2} y_{1}+\epsilon y_{2}+\cdots
\end{aligned}
$$

As before:

$$
\begin{aligned}
& y=0+\epsilon^{1 / 2}+O(\epsilon) \\
& y= \pm 1-\frac{1}{2} \epsilon^{1 / 2}+O(\epsilon) \\
& x=1+\epsilon+3 \epsilon^{2}+\cdots \\
& x=1+O\left(\epsilon^{1 / 2}\right) \\
& x= \pm \frac{1}{\epsilon^{1 / 2}}-\frac{1}{2}+O\left(\epsilon^{1 / 2}\right)
\end{aligned}
$$

## Example 6.2.

$$
\begin{aligned}
(1-\epsilon) x^{2}-2 x+1 & =0 \\
x & =x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}+\cdots \\
x^{2} & =x_{0}^{2}+2 \epsilon x_{0} x_{1}+\epsilon^{2}\left(2 x_{0} x_{2}+x_{1}^{2}\right)+\cdots \\
(1-\epsilon)\left[x_{0}^{2}+2 \epsilon x_{0} x_{1}+\epsilon^{2}\left(2 x_{0} x_{2}+x_{1}^{2}\right)+\cdots\right]-2\left(x_{0}+\epsilon x_{1}+\epsilon^{2} x_{2}\right)+1 & =O\left(\epsilon^{3}\right) \\
x_{0}^{2}-2 x_{0}+1 & =0 \\
2 x_{0} x_{1}-x_{0}^{2}-2 x_{1} & =0 \\
2\left(x_{0}-1\right) x_{1} & =x_{0}^{2} \\
x_{0} & =1
\end{aligned}
$$

There is no solution of the assumed form (perturbing off a repeated root).

$$
x=1 \pm \sqrt{\epsilon}
$$

The correct expansion is

$$
x=x_{0}+\epsilon^{1 / 2} x_{1}+\epsilon x_{2}+\cdots
$$

## 7 4-16-12

### 7.1 Asymptotic and Convergent Series

Euler 1754:

$$
I(x)=\int_{0}^{\infty} \frac{e^{-t}}{1+x t} d t
$$

How does $I(x)$ behave as $x \rightarrow 0^{+}$? This integral is well-defined for $x \geq 0$.
Formally: for small $x$,

$$
\begin{align*}
\frac{1}{1+x t} & =1-x t+(x t)^{2}-\cdots+(-1)^{n}(x t)^{n}+\cdots \\
I(x) & =\int_{0}^{\infty} e^{-t} d t-x \int_{0}^{\infty} t e^{-t} d t+\cdots+(-1)^{n} x^{n} \int_{0}^{\infty} t^{n} e^{-t} d t+\cdots \\
& =1-x+2 x^{2}+\cdots+(-1)^{n} n!x^{n}+\cdots \\
I(x) & =\sum_{n=0}^{\infty}(-1)^{n} n!x^{n} \tag{7.1}
\end{align*}
$$

For example, at $x=1$ :

$$
\int_{0}^{\infty} \frac{e^{-t}}{1+t} d t=1-2!+3!-4!+5!\cdots
$$

The ratio test shows that (7.1) has zero radius of convergence, so it diverges for all $x \neq 0$. Where did we go wrong? The expansion for $\frac{1}{1+x t}$ is only valid for $x t<1$. So our expansion doesn't converge everywhere, namely when $t$ is large. But when $t$ is large, we have exponential decay in our integral.

For example, at $x=0.1$ :

$$
\begin{aligned}
\sum_{n=0}^{12}(-1)^{n} n!x^{n} & =0.91542 \\
\int_{0}^{\infty} \frac{e^{-t}}{1+(0.1) t} d t & =0.9156
\end{aligned}
$$

## Theorem 7.1.

$$
\begin{aligned}
& x \geq 0, N=0,1,2, \ldots . \\
& \quad\left|I(x)-\sum_{n=0}^{N}(-1)^{n} n!x^{n}\right| \leq(N+1)!x^{N+1}
\end{aligned}
$$

Proof.

$$
\begin{aligned}
I(x) & =\int_{0}^{\infty} \frac{e^{-t}}{1+x t} d t \\
& =1-\int_{0}^{\infty} \frac{e^{-t}}{(1+x t)^{2}} d t \\
& =1-x+\cdots+(-1)^{N} N!x^{N}+R_{N+1}(x) \\
R_{N+1}(x) & =(-1)^{N+1}(N+1)!x^{N+1} \int_{0}^{\infty} \frac{e^{-t}}{(1+x t)^{N+2}} d t \\
\left|R_{N+1}(x)\right| & \leq(N+1)!x^{N+1} \underbrace{\int_{0}^{\infty} e^{-t} d t}_{=1}
\end{aligned}
$$

We write this as

$$
I(x)=\sum_{n=0}^{N}(-1)^{n} n!x^{n}+O\left(x^{N+1}\right) \quad \text { as } x \rightarrow 0^{+}
$$

$O\left(x^{N+1}\right)$ stands for a term bounded by a constant times $|x|^{N+1}$.
Convergent: Fix $x, N \rightarrow \infty$
Asymptotic: Fix $N, x \rightarrow 0^{+}$

### 7.1.1 Optimal Truncation

$$
|I(x)-\underbrace{\sum_{n=0}^{N}(-1)^{n} n!x^{n}}_{S_{N}(x)}| \leq(N+1)!x^{N+1}
$$

As long as the $x$ power is beating out the factorial, the error is going down. The optimal truncation is at $N \sim\left[\frac{1}{x}\right]$. Then the error is

$$
\begin{aligned}
\text { Error } & \sim\left(\frac{1}{x}\right)!x^{1 / x} \\
& \sim \sqrt{\frac{2 \pi}{x}} e^{-1 / x} \quad \text { as } x \rightarrow 0^{+}
\end{aligned}
$$

where we have used Stirling's formula:

$$
n!\sim \sqrt{2 \pi} n^{n+\frac{1}{2}} e^{-n} \quad \text { as } n \rightarrow \infty
$$

So we get exponential accuracy by optimal truncation (asymptotics beyond all orders).

### 7.2 Notation for Asymptotic Behavior

$f(x), g(x), x \rightarrow x_{0}\left(x_{0}=0^{+}, \infty, \ldots\right)$
We write $f(x)=O(g(x))$ as $x \rightarrow x_{0}$ if there exist constants $C, \delta>0$ such that

$$
|f(x)| \leq C|g(x)| \quad \text { for } \quad\left|x-x_{0}\right|<\delta
$$

We write that $f(x)=o(g(x))$ if for all $\epsilon>0$ there exists $\delta>0$ such that

$$
|f(x)| \leq \epsilon|g(x)| \quad \text { for } \quad\left|x-x_{0}\right|<\delta .
$$

If $g(x) \neq 0$, this is equivalent to

$$
\lim _{x \rightarrow x_{0}}\left|\frac{f(x)}{g(x)}\right|=0
$$

$o$ implies $O$.
Example 7.2.

$$
\begin{aligned}
& f(x)=x \\
& g(x)=x^{2}
\end{aligned}
$$

As $x \rightarrow 0, x^{2}=o(x)$. As $x \rightarrow \infty, x=o\left(x^{2}\right)$.

$$
\begin{aligned}
& f(x)=\sin \left(\frac{1}{x}\right) \\
& g(x)=x
\end{aligned}
$$

As $x \rightarrow 0$, there is no relation between $f$ and $g$. But we can say that $\sin \left(\frac{1}{x}\right)=O(1)$ as $x \rightarrow 0$.

$$
\begin{aligned}
f(x) & =x \\
g(x) & =10^{6} \log x
\end{aligned}
$$

As $x \rightarrow \infty, 10^{6} \log x=o(x)$. Similarly, $10^{6} \log (\log x)=o(\log x)$ as $x \rightarrow \infty$.

$$
\begin{aligned}
& f(x)=x \\
& g(x)=\log \frac{1}{x}
\end{aligned}
$$

As $x \rightarrow 0, x=o\left(\frac{1}{\log \frac{1}{x}}\right)$.
$e^{-1 / x}=o\left(x^{n}\right)$ as $x \rightarrow 0^{+}$.

### 8.1 Perturbation Theory for ODE's

1. Regular perturbation problems
2. Singular perturbation problems
(a) Boundary/initial layer problems. These are treated by the method of matched asymptotic expansions (MMAE)
(b) Oscillation problems. These are treated by the method of multiple scales (MMS)

### 8.2 Overdamped Simple Harmonic Oscillator (Logan 2.4)

$$
\begin{aligned}
m \ddot{y}+a \dot{y}+k y & =0 \\
y(0) & =0 \\
\dot{y}(0) & =\frac{I}{m}
\end{aligned}
$$

Dimensions: mass $M$, length $L$, and time $T$
Parameters: $m(M), a\left(\frac{M}{T}\right), k\left(\frac{M}{T^{2}}\right), I\left(\frac{M L}{T}\right)$
Variables: $y(L), t(T)$
For large damping, choose time scale $\frac{a}{k}$ (which has dimension $T$ ). Choose length scale $\frac{I}{a}$ (which has dimension L). Set

$$
\begin{aligned}
y & =\frac{I}{a} y^{*} \\
t & =\frac{a}{k} t^{*} \\
\frac{d}{d t} & =\frac{k}{a} \frac{d}{d t^{*}}
\end{aligned}
$$

(Henceforth, dots will denote derivatives with respect to $t^{*}$.) Since the equation is linear, the rescaling factor of $y$ will cancel out. So we have

$$
\begin{aligned}
m\left(\frac{k}{a}\right)^{2} \ddot{y}^{*}+a\left(\frac{k}{a}\right) \dot{y}^{*}+k y^{*} & =0 \\
y^{*}(0) & =0 \\
\left(\frac{k}{a}\right)\left(\frac{I}{a}\right) \dot{y}^{*}(0) & =\frac{I}{m} \\
\frac{m k}{a^{2}} \ddot{y}^{*}+\dot{y}^{*}+y^{*} & =0 \\
y^{*}(0) & =0 \\
\dot{y}^{*}(0) & =\frac{a^{2}}{m k} \\
\epsilon: & =\frac{m k}{a^{2}}
\end{aligned}
$$

Nondimensionalized problem (drop the *'s):

$$
\begin{aligned}
\epsilon \ddot{y}+\dot{y}+y & =0 \\
y(0) & =0 \\
\dot{y}(0) & =\frac{1}{\epsilon}
\end{aligned}
$$

We want to find the approximate solution when $\epsilon$ is small (and positive). This is a singular perturbation problem because if we set $\epsilon=0$ then we change the order of the ODE from 2nd order to 1st order. We can't solve a 1st order ODE with 2 initial conditions.

The solution consists of two parts:
(a) a short initial layer where $\ddot{y}$ is large $\Rightarrow$ fast
(b) long outer regions where $\ddot{y}$ is $O(1) \Rightarrow$ slow

Idea: construct different "inner" and "outer" approximations, then match them.
Outer solution (b)

$$
\begin{aligned}
y & =y_{0}(t)+\epsilon y_{1}(t)+\epsilon^{2} y_{2}(t) \ldots \\
\epsilon \ddot{y}_{0}+\epsilon^{2} \ddot{y}_{1}+\dot{y}_{0}+\epsilon \dot{y}_{1}+\epsilon^{2} \dot{y}_{2}+y_{0}+\epsilon y_{1}+\epsilon^{2} y_{2} & =O\left(\epsilon^{3}\right) \\
\dot{y}_{0}+y_{0} & =0 \\
\ddot{y}_{0} \dot{y}_{1}+y_{1} & =0 \\
\dot{y}_{n}+y_{n}+\ddot{y}_{n-1} & =0 \\
y_{0}(t) & =c e^{-t}, \quad t=O(1)
\end{aligned}
$$

This is the leading order outer solution.
Initial layer (a)
Say $t=O(\delta)$. Introduce the time variable

$$
\begin{aligned}
T & =\frac{t}{\delta} \\
\frac{d}{d t} & =\frac{1}{\delta} \frac{d}{d t} \\
y(t ; \epsilon) & =Y(T ; \epsilon) \\
\frac{\epsilon}{\delta^{2}} \frac{d^{2} Y}{d T^{2}}+\frac{1}{\delta} \frac{d Y}{d T}+Y & =0
\end{aligned}
$$

The dominant balances will be

1. $\frac{1}{\delta}=1, \delta=1$ (outer)
2. $\frac{\epsilon}{\delta^{2}}=\frac{1}{\delta}, \delta=t$ (inner)
3. The third possibility, $\frac{\epsilon}{\delta^{2}}=1$, is not a dominant balance

We get

$$
\begin{aligned}
\frac{d^{2} Y}{d T^{2}}+\frac{d Y}{d T}+\epsilon Y & =0 \\
Y(0) & =0 \\
\frac{d Y}{d T}(0) & =1
\end{aligned}
$$

So the inner expansion is:

$$
\begin{aligned}
Y & =Y_{0}(T)+\epsilon Y_{1}(T)+O\left(\epsilon^{2}\right) \\
\frac{d^{2} Y_{0}}{d T^{2}}+\frac{d Y_{0}}{d T} & =0 \\
Y_{0}(0) & =0 \\
\frac{d Y_{0}}{d T}(0) & =1 \\
Y_{0}(T) & =A+B e^{-T}=1-e^{-T}
\end{aligned}
$$

The leading order inner solution is

$$
\begin{aligned}
Y_{0}(T) & =1-e^{-T} \\
T & =\frac{t}{\epsilon}=O(1)
\end{aligned}
$$



The matching condition is

$$
\begin{aligned}
\lim _{T \rightarrow \infty} Y_{0}(T) & =\lim _{t \rightarrow 0^{+}} y_{0}(t) \\
1 & =C \\
y(t, \epsilon) & \sim\left\{\begin{array}{rr}
1-e^{-t / \epsilon} & t=O(\epsilon) \\
e^{-t} & t=O(1)
\end{array}\right.
\end{aligned}
$$

## $9 \quad 4-20-12$

### 9.1 Strongly Damped Oscillator

Remark 9.1. A note on expansions

$$
\begin{aligned}
(1+x)^{\alpha} & =1+\alpha x+\frac{1}{2} \alpha(\alpha-1) x^{2}+\frac{1}{3!} \alpha(\alpha-1)(\alpha-2) x^{3}+\cdots, \quad|x|<1 \\
\sqrt{1+x} & =1+\frac{1}{2} x+\frac{1}{2}\left(\frac{1}{2}\right)\left(-\frac{1}{2}\right) x^{2}+\cdots \\
& =1+\frac{1}{2} x-\frac{1}{8} x^{2}+\cdots \\
\frac{1}{1+x} & =1-x+x^{2}-x^{3}+\cdots
\end{aligned}
$$

$$
\begin{aligned}
\epsilon \ddot{y}+\dot{y}+y & =0 \\
y(0) & =0 \\
\dot{y}(0) & =\frac{1}{\epsilon}
\end{aligned}
$$

The characterisitic equation, $y=e^{r t}$, gives

$$
\begin{aligned}
\epsilon r^{2}+r+1 & =0 \\
r_{ \pm} & =\frac{-1 \pm \sqrt{1-4 \epsilon}}{2 \epsilon} \\
r_{-} & =-\frac{1}{\epsilon}+O(1) \\
r_{+} & =\frac{-1+\left(1-\frac{1}{2} \cdot 4 \epsilon+O(\epsilon)^{2}\right)}{2 \epsilon} \\
& =-1+O(\epsilon) \\
y(t) & =A e^{r_{-} t}+B e^{r_{+} t} \\
y(0) & =0 \quad A+B=0 \\
\dot{y}(0) & =\frac{1}{\epsilon} \quad r_{-} A+r_{+} B=\frac{1}{\epsilon} \\
B & =-A \\
A & =\frac{1}{\epsilon}\left(\frac{1}{r_{-}-r_{+}}\right) \\
B & =\frac{1}{\epsilon}\left(\frac{1}{r_{+}-r_{-}}\right) \\
r_{+}-r_{-} & =\frac{-1+\sqrt{1-4 \epsilon}}{2 \epsilon}-\left(\frac{-1-\sqrt{1-4 \epsilon}}{2 \epsilon}\right) \\
& =\frac{\sqrt{1-4 \epsilon}}{\epsilon}
\end{aligned}
$$

$$
\text { Exact solution: } \quad y(t)=-\frac{1}{\sqrt{1-4 \epsilon}} \exp \left[-\frac{(1+\sqrt{1-4 \epsilon})}{2 \epsilon} t\right]
$$

$$
+\frac{1}{\sqrt{1-4 \epsilon}} \exp \left[-\frac{(1-\sqrt{1-4 \epsilon})}{2 \epsilon} t\right]
$$

As $\epsilon \rightarrow 0^{+}$,

$$
\begin{aligned}
y(t) & \sim-e^{-t / \epsilon}+e^{t} \\
t & =\epsilon T \\
y & =-e^{-T}+e^{\epsilon T}
\end{aligned}
$$

Balancing $\epsilon \ddot{y}+\dot{y}$ gives $e^{-t / \epsilon}$, while balancing $\dot{y}+y$ gives $e^{-t}$.
As $\epsilon \rightarrow 0^{+}$,

$$
y(t) \sim\left\{\begin{array}{rl}
1-e^{-t / \epsilon} & t=O(\epsilon) \\
e^{t} & t=O(1), t>0
\end{array}\right.
$$



### 9.2 Phase Plane

$$
\begin{aligned}
\epsilon \ddot{y}+\dot{y}+y & =0 \\
\dot{y} & =z \\
\dot{z} & =-\frac{1}{\epsilon}(y+z)
\end{aligned}
$$

Two regimes:

1. "Slow" manifold, $y+z=0$. The approximate equation for $y$ is then

$$
\dot{y}=-y \quad \Rightarrow \quad y=c e^{-t}
$$

2. "Fast" system, $\dot{z}=O(1 / \epsilon)$ and $\dot{y}=O(1)$.

$$
\begin{aligned}
T & =\frac{t}{\epsilon} \\
\frac{d}{d t} & =\frac{1}{\epsilon} \frac{d}{d T} \\
\frac{1}{\epsilon} \frac{d y}{d T} & =z \\
\frac{1}{\epsilon} \frac{d z}{d T} & =-\frac{1}{\epsilon}(y+z) \\
\frac{d y}{d T} & =\epsilon z \approx 0 \\
\frac{d z}{d T} & =-(y+z)
\end{aligned}
$$

$y+z \neq 0$, so the approximate equation is

$$
\begin{aligned}
\dot{y} & =0 \\
\dot{z} & =-\frac{1}{\epsilon}(z+y)
\end{aligned}
$$



Figure 3: "Geometric Singular Perturbation Theory"

### 9.3 Michaelis Menton Enzyme Kinetics

$$
\begin{aligned}
& \mathrm{H}_{2} \mathrm{O}_{2} \rightarrow \mathrm{H}_{2} \mathrm{O}+\mathrm{O} \\
& \mathrm{E}+\mathrm{S} \stackrel{k_{0}}{\longleftrightarrow} \xrightarrow{k_{1}} \mathrm{C} \xrightarrow{k_{2}} \mathrm{P}
\end{aligned}
$$

Law of mass actions:
rate of reaction $\propto$ product of concentrations,
where the constant of proportionality is the rate constant.

- $e(t)=$ concentration of E
- $s(t)=$ concentration of S
- $c(t)=$ concentration of C
- $p(t)=$ concentration of P

$$
\begin{aligned}
& \frac{d e}{d t}=-k_{1} e_{s}+\left(k_{0}+k_{2}\right) c \\
& \frac{d s}{d t}=-k_{1} e s+k_{0} c \\
& \frac{d c}{d t}=k_{1} e s-\left(k_{0}+k_{2}\right) c \\
& \frac{d p}{d t}=k_{2} c
\end{aligned}
$$

We see that

$$
\begin{aligned}
\frac{d}{d t}(e+c) & =0 \\
e+c & =\mathrm{constant}
\end{aligned}
$$

### 10.1 Enzyme Kinetics (Continued)

$$
\begin{aligned}
\mathrm{E}+\mathrm{S} \stackrel{k_{0}}{ } & \xrightarrow{k_{1}} \mathrm{C} \xrightarrow{k_{2}} \mathrm{P} \\
\frac{d e}{d t} & =-k_{1} e_{s}+\left(k_{0}+k_{2}\right) c \\
\frac{d s}{d t} & =-k_{1} e s+k_{0} c \\
\frac{d c}{d t} & =k_{1} e s-\left(k_{0}+k_{2}\right) c \\
\frac{d p}{d t} & =k_{2} c \\
e(0) & =e_{0} \\
s(0) & =s_{0} \\
c(0) & =0 \\
p(0) & =0 \\
e+c & =e_{0} \\
\frac{d}{d t}[e+c] & =0 \\
\frac{d e}{d t} & =-k_{1} e s+\left(k_{0}+k_{2}\right)\left(e_{0}-e\right) \\
\frac{d s}{d t} & =-k_{1} e s+k_{0}\left(e_{0}-e\right)
\end{aligned}
$$

Dimensions: time $T$, concentration $C$
Independent Variables: $t(T)$
Dependent Variables: $e(C), s(C)$
Parameters: $e_{0}(C), s_{0}(C), k_{0}\left(\frac{1}{T}\right), k_{1}\left(\frac{1}{C T}\right), k_{2}\left(\frac{1}{T}\right)$

$$
\begin{aligned}
u(\tau) & =\frac{s(t)}{s_{0}} \\
v(\tau) & =\frac{c(t)}{e_{0}} \\
\tau & =k_{1} e_{0} t \\
\frac{d u}{d \tau} & =-u+(u+k-\lambda) v \\
\epsilon \frac{d v}{d \tau} & =u-(u+k) v \\
u(0) & =1 \\
v(0) & =0 \\
\epsilon & =\frac{e_{0}}{s_{0}} \\
k & =\frac{k_{0}+k_{2}}{k_{1} s_{0}} \\
\lambda & =\frac{k_{2}}{k_{1} s_{0}}
\end{aligned}
$$

We have two regimes:
(a) Short time, $\tau=O(\epsilon)$
(b) Long time, $\tau=O(1)$
(b) Long time. Expand

$$
\begin{aligned}
u & =u_{0}(\tau)+\epsilon u_{1}(\tau)+\cdots \\
v & =v_{0}(\tau)+\epsilon v_{1}(\tau)+\cdots \\
\frac{d u_{0}}{d \tau} & =-u_{0}+\left(u_{0}+k-\lambda\right) v_{0} \\
0 & =u_{0}-\left(u_{0}+k\right) v_{0} \\
v_{0} & =\frac{u_{0}}{u_{0}+k} \\
\frac{d u_{0}}{d \tau} & =-u_{0}+\left(u_{0}+k-\lambda\right) \cdot \frac{u_{0}}{u_{0}+k} \\
& =-\frac{\lambda u_{0}}{u_{0}+k}
\end{aligned}
$$

(a) Short time.

$$
\begin{aligned}
T & =\frac{\tau}{\epsilon} \\
\frac{d}{d t} & =\frac{1}{\epsilon} \frac{d}{d T} \\
U(T) & =u(t) \\
\frac{d U}{d T} & =\epsilon[-U+(U+k-\lambda) V] \\
\frac{d V}{d T} & =U-(U+k) V \\
U & =U_{0}+\epsilon U_{1}+\cdots \\
V & =V_{0}+\epsilon V_{1}+\cdots \\
\frac{d U_{0}}{d T} & =0 \\
\frac{d V_{0}}{d T} & =U_{0}-\left(U_{0}+k\right) V_{0} \\
U_{0}(0) & =1 \\
V_{0}(0) & =0 \\
U_{0}(T) & =1 \\
\frac{d V_{0}}{d T} & =1-(1+k) V_{0} \\
V_{0}(0) & =0 \\
V_{0}(T) & =\frac{1-e^{-(1+k) T}}{1+k}
\end{aligned}
$$

(b) Matching.

$$
u_{0}(0)=\lim _{T \rightarrow \infty} U_{0}(T)=1
$$



Figure 4: $\mathrm{E}+\mathrm{S} \stackrel{k_{0}}{\longleftrightarrow} \mathrm{C} \xrightarrow{k_{1}} \mathrm{P}$

### 11.1 Geometric Singular Perturbation Theory

$$
\begin{aligned}
\epsilon \dot{x} & =f(x, y) \\
\dot{y} & =g(x, y)
\end{aligned}
$$

$x(t) \in \mathbb{R}^{m}, y(t) \in \mathbb{R}^{n}, f: \mathbb{R}^{m} \times \mathbb{R}^{n} \rightarrow \mathbb{R}^{m}, g: \mathbb{R}^{m} \times \mathbb{R}^{n} \rightarrow \mathbb{R}^{n} . x$ contains the "fast" variables, $y$ contains the "slow" variables. Introduce a fast time: $T=\frac{t}{\epsilon}$. Let ${ }^{\prime}=\frac{d}{d T}$ and ${ }^{*}=\frac{d}{d t}$. So $\frac{1}{\epsilon} \frac{d}{d T}=\frac{d}{d t}$.

$$
\begin{aligned}
x^{\prime} & =f(x, y) \\
y^{\prime} & =\epsilon g(x, y)
\end{aligned}
$$

"Slow" system:

$$
\begin{aligned}
f(x, y) & =0 \\
\dot{y} & =g(x, y)
\end{aligned}
$$

"Fast" system:

$$
\begin{aligned}
x^{\prime} & =f(x, y) \\
y^{\prime} & =0
\end{aligned}
$$

The slow manifold is $f(x, y)=0$. We can't satisfy all of the initial data in the slow system, because the initial data for $x$ has to be such that $f(x, y)=0$. Physicists say that the $x$ variable is a slave to the $y$ variable.

For the fast system, $y=y_{0}$ (constant) and $x^{\prime}=f\left(x, y_{0}\right)$.
Simplest case:

- The slow manifold is a graph, $x=\phi(y), \phi: \mathbb{R}^{n} \rightarrow \mathbb{R}^{m}$.


Figure 5: $f(\phi(y), y)=0, \dot{y}=g(\phi(y), y)$.

- Assume that $x=\phi(y)$ is a globally asymptotically stable (unique) equilibrium for the "fast" equation, $x^{\prime}=f(x, y)$.

Tikhonov (1948) and Levinson (1949) gave a theory for attracting slow manifolds in these "fast-slow" systems.

Fenichel (1971) proved that the full system has an invariant manifold close to the slow manifold for small $\epsilon$ provided $x=\phi(y)$ is a hyperbolic equilibrium of the "fast" system $x^{\prime}=f(x, y)$.


### 11.2 Van der Pol Oscillator

$$
\underbrace{\epsilon \ddot{x}+\left(x^{2}-1\right) \dot{x}}_{=-\dot{y}}+x=0
$$

Small mass/large damping: $0<\epsilon \ll 1$
Negative damping/excitability: $|x|<1$
Positive damping: $|x|>1$
Lienard variables:

$$
\begin{aligned}
y & =x-\frac{1}{3} \dot{x}^{3}-\epsilon \dot{x} \\
\epsilon \dot{x} & =x-\frac{1}{3} x^{3}-y \\
\dot{y} & =x
\end{aligned}
$$

Slow manifold: $y=x-\frac{1}{3} x^{3}$


Figure 6:

Slow system

$$
\begin{aligned}
& y=x-\frac{1}{3} x^{3} \\
& \dot{y}=x
\end{aligned}
$$

Fast system

$$
\begin{aligned}
& x^{\prime}=x-\frac{1}{3} x^{3}-y \\
& y^{\prime}=0
\end{aligned}
$$

## $12 \quad 4-27-12$

12.1 Heat Flow in a Slowly-Varying Rod


Figure 7: $u(x, t)=$ temperature

$$
\begin{aligned}
u_{t} & =\nu u_{x x}, \quad 0<x<L(t), t>0 \\
u(0, t) & =0 \\
u(L(t), t) & =g(t) \\
u(x, 0) & =f(x)
\end{aligned}
$$

$$
\begin{aligned}
L_{0} & =L(0) \\
T_{0} & =\text { time-scale of variations in } L(t) \\
\theta & =\text { typical temperature } \\
L(t) & =L_{0} L^{*}\left(\frac{t}{T_{0}}\right) \\
g(t) & =\theta_{0} g^{*}\left(\frac{t}{T_{0}}\right) \\
f(x) & =\theta_{0} f^{*}\left(\frac{x}{L_{0}}\right) \\
x^{*} & =\frac{x}{L_{0}} \\
t^{*} & =\frac{t}{T_{0}} \\
u^{*} & =\frac{u}{\theta_{0}} \\
\partial_{x} & =\frac{1}{L_{0}} \partial_{x^{*}} \\
\partial_{t} & =\frac{1}{T_{0}} \partial_{t^{*}} \\
u_{t} & =\frac{\theta_{0}}{T_{0}} u_{t^{*}}^{*} \\
u_{x x} & =\frac{\theta_{0}}{L_{0}^{2}} u_{x^{*} x^{*}}^{*} \\
u_{t} & =\nu u_{x x} \\
\frac{\theta_{0}}{T_{0}} u_{t^{*}}^{*} & =\frac{\nu \theta_{0}}{L_{0}^{2}} u_{x^{*} x^{*}}^{*} \\
\epsilon u_{t^{*}}^{*} & =u_{x^{*} x^{*}}^{*} \\
\epsilon & =\frac{L_{0}^{2}}{\nu T_{0}}
\end{aligned}
$$

So we have

$$
\begin{aligned}
\epsilon u_{t^{*}}^{*} & =u_{x^{*} x^{*}}^{*}, \quad 0<x^{*}<L^{*}\left(t^{*}\right), t^{*}>0 \\
u^{*}\left(0, t^{*}\right) & =0 \\
u^{*}\left(L^{*}\left(t^{*}\right), t^{*}\right) & =g^{*}\left(t^{*}\right) \\
u^{*}\left(x^{*}, 0\right) & -f^{*}\left(x^{*}\right)
\end{aligned}
$$

Interpretation of $\epsilon$ :

- $T_{d}=$ diffusion-timescale, i.e. time, for heat to diffuse from one end of the rod to the other. $L \sim$ $\sqrt{\nu T} \Leftrightarrow T \sim L^{2} / \nu$.
- $T_{d}=\frac{L_{0}^{2}}{\nu}$
- $\epsilon=\frac{T_{d}}{T_{0}}$

Assume $\epsilon \ll 1$. This means that heat diffuses rapidly over the rod relative to the timescale of variations in the length/boundary data.

Drop the *'s.

$$
\begin{array}{rlrl}
\epsilon u_{t} & =u_{x x}, & & 0<x<L(t), t>0 \\
u(0, t) & =0 & & \\
u(L(t), t) & =g(t) & & \\
u(x, 0) & =f(x), \quad 0<x<1, L(0)=1
\end{array}
$$

Outer expansion:

$$
\begin{aligned}
u & =u_{0}(x, t)+\epsilon u_{1}(x, t)+O\left(\epsilon^{2}\right) \\
u_{0, x x} & =0, \quad 0<x<L \\
u_{0}(0, t) & =0 \\
u_{0}(L, t) & =g
\end{aligned}
$$

We have to drop the initial condition (because we wouldn't be able to satisfy it with the outer solution).

$$
\begin{aligned}
u_{0}(x, t) & =A(t) x+B(t) \\
& =\frac{g(t)}{L(t)} x
\end{aligned}
$$

Inner expansion:

$$
\begin{aligned}
T & =\frac{t}{\epsilon} \\
u(x, t ; \epsilon) & =U(x, T ; \epsilon) \\
\partial_{t} & =\frac{1}{\epsilon} \partial_{T} \\
U_{t} & =U_{x x}, \quad 0<x<L(\epsilon T), T>0 \\
U(0, T) & =0 \\
U(L(\epsilon T), \epsilon T) & =g(\epsilon T) \\
U(x, 0) & =f(x), \quad 0<x<1 \\
U & =U_{0}(x, T)+\epsilon U_{1}(x, T)+O\left(\epsilon^{2}\right) \\
U_{0, T} & =U_{0, x x}, \quad 0<x<1, T>0 \\
U_{0}(0, T) & =0 \\
U_{0}(1, T) & =g(0) \\
U_{0}(x, 0) & =f(x), \quad 0<x<1
\end{aligned}
$$

Solve by separating variables.

$$
\begin{aligned}
U(x, T) & =g(0) X+V(x, T) \\
V_{t} & =V_{x x} \\
V(0, T) & =0 \\
V(1, T) & =0 \\
V(x, 0) & =f(x)-g(0) x \\
V(x, T) & =\sum_{n=1}^{\infty} c_{n} e^{-n^{2} \pi^{2} T} \sin (n \pi x) \\
c_{n} & =2 \int_{0}^{1}[f(x)-g(0) x] \sin (n \pi x) d x \\
U_{0}(x, T) & =g(0) x+V(x, T)
\end{aligned}
$$

So we have

$$
\begin{aligned}
\text { Outer solution: } \quad u_{0}(x, t) & =\frac{g(t)}{L(t)} x \\
\text { Inner solution: } \quad U_{0}(x, T) & =g(0) x+V(x, T)
\end{aligned}
$$

Do they match?

$$
\begin{aligned}
\lim _{T \rightarrow \infty} U_{0}(x, T) & =g(0) x \\
\lim _{t \rightarrow 0^{+}} u_{0}(x, t) & =g(0) x
\end{aligned}
$$

Uniform solution:

$$
\begin{aligned}
u & \sim u_{\text {inner }}+u_{\text {outer }}-u_{\text {matching }} \\
& \sim \frac{g(t)}{L(t)} x+V\left(x, \frac{t}{\epsilon}\right)
\end{aligned}
$$

## $13 \quad 4-30-12$

### 13.1 Boundary Layer Problems

Navier-Stokes equation for incompressible fluid:

$$
\begin{array}{rlrl}
\vec{u}_{t} \vec{u} \cdot \nabla \vec{u}+\nabla p & =\epsilon \Delta \vec{u}, & \epsilon=\frac{1}{\mathrm{Re}} \\
\nabla \cdot \vec{u} & =0 \quad & \text { ("no slip" condition) } \\
\vec{u}(\vec{x}, 0) & =\vec{u}_{0}(\vec{x}) & \\
\vec{u}(\vec{x}, t) & =0 \quad \text { on } \partial \Omega
\end{array}
$$

Setting $\epsilon=0$ (no viscosity), we get the Euler equation:

$$
\vec{u}_{t}+\vec{u} \cdot \nabla \vec{u}+\nabla p=0
$$

The Euler equation with no-slip boundary condition is overdetermined. So we impose the "no-flow" condition:

$$
\vec{u} \cdot \vec{n}=0
$$

Prandtl (1905) introduced boundary layer theory.


The velocity goes quickly from zero to something large, so the derivative is very large.

### 13.2 Model Boundary Layer Problem

$$
\begin{aligned}
\epsilon y^{\prime \prime}+2 y^{\prime}+y & =0, \quad 0<x<1 \\
y(0) & =0 \\
y(1) & =1
\end{aligned}
$$

We want to find an asymptotic approximation of the solution for $0<\epsilon \ll 1$.
Straightforward (outer) expansion:

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\epsilon^{2} y_{2}(x)+O\left(\epsilon^{3}\right) \\
2 y_{0}^{\prime}+y_{0} & =0 \\
2 y_{1}^{\prime}+y_{1}+y_{0}^{\prime \prime} & =0 \\
2 y_{n}^{\prime}+y_{n}+y_{n-1}^{\prime \prime} & =0
\end{aligned}
$$

Problem: can't satisfy both BC's because the order of the ODE drops from 2 to 1 at $\epsilon=0$. It turns out that the correct BC to impose is the BC at $x=1$.

$$
\begin{aligned}
y_{0}(1) & =1 \\
y_{1}(1) & =0 \\
y_{n}(1) & =0 \\
y_{0}(x) & =c e^{-x / 2} \\
& =e^{1 / 2} e^{-x / 2}
\end{aligned}
$$

So we get a boundary layer near $x=0$ where the solution adjusts rapidly from $\approx e^{1 / 2}$ to 0 at $x=0$.
Inner expansion (near $x=0$ ):

$$
\begin{aligned}
X & =\frac{x}{\delta} \\
y(x ; \epsilon) & =Y(X ; \epsilon) \\
y^{\prime}(x ; \epsilon) & =\frac{1}{\delta} \frac{d Y}{d X}=\frac{1}{\delta} Y^{\prime} \\
\underbrace{\frac{\epsilon}{\delta^{2}}}_{(1)}+\underbrace{\frac{2}{\delta} Y^{\prime}}_{(2)}+\underbrace{Y}_{3} & =0
\end{aligned}
$$

Dominant balances:

- (1) ~ (2): $\frac{\epsilon}{\delta^{2}}=\frac{1}{\delta} \Rightarrow \delta=\epsilon,(3 \lll(1) \sim(2)$
- (2) ~ (3): $\delta=1 \Rightarrow(1) \ll(2) \sim(3)$
- (1) ~ (3): $\frac{\epsilon}{\delta^{2}}=1 \Rightarrow \delta=\epsilon^{1 / 2}$, (2) $\gg(1) \sim(3)$

Take $\delta=\epsilon$.

$$
\begin{aligned}
Y^{\prime \prime}+2 Y^{\prime}+\epsilon Y & =0 \\
Y & =Y_{0}(X)+\epsilon Y_{1}(X)+\cdots \\
Y_{0}^{\prime \prime}+2 Y_{0}^{\prime} & =0 \\
Y_{1}^{\prime \prime}+2 Y_{1}^{\prime}+Y_{0} & =0 \\
Y_{0}(0) & =0 \\
Y_{0}^{\prime} & =c e^{-2 X} \\
Y_{0}(X) & =c_{1}+c_{2} e^{-2 X}=c\left(1-e^{-2 X}\right)
\end{aligned}
$$

Matching condition:

$$
\begin{aligned}
\lim _{x \rightarrow 0^{+}} y_{0}(x) & =\lim _{X \rightarrow \infty} Y_{0}(X) \\
e^{1 / 2} & =c \\
Y_{0}(X) & =e^{1 / 2}\left(1-e^{-2 X}\right)
\end{aligned}
$$

Leading-order asymptotic solution:

$$
y(x ; \epsilon) \sim\left\{\begin{aligned}
e^{1 / 2} e^{-x / 2} & \text { as } \epsilon \rightarrow 0^{+}, 0<x \leq 1 \\
e^{1 / 2}\left(1-e^{-2 x / \epsilon}\right) & 0 \leq \frac{x}{\epsilon}<\infty
\end{aligned}\right.
$$



Uniform solution:

$$
\begin{aligned}
& y_{\text {inner }}+y_{\text {outer }}-y_{\text {overlap }} \\
& y(x ; \epsilon) \sim e^{1 / 2}\left(e^{-x / 2}-e^{-2 x / \epsilon}\right)
\end{aligned}
$$

Let's compare this to the exact solution. The characteristic equation is

$$
\begin{aligned}
\epsilon r^{2}+2 r+1 & =0 \\
r & =\frac{-1 \pm \sqrt{1-\epsilon}}{\epsilon} \\
r & =-\alpha(\epsilon),-\frac{\beta(\epsilon)}{\epsilon} \\
\beta(\epsilon) & =2+\cdots \\
-1+\sqrt{1-\epsilon} & =-1+\left(1-\frac{1}{2} \epsilon\right)=-\frac{1}{2} \epsilon \\
y(x ; \epsilon) & =\frac{e^{-\alpha x}-e^{-\beta x / \epsilon}}{e^{-\alpha}-e^{-\beta / \epsilon}} \\
& \sim \frac{e^{-x / 2}-e^{-2 x / \epsilon}}{e^{-1 / 2}-e^{-2 / \epsilon}}
\end{aligned}
$$

This agrees with the uniform solution (to leading order in $\epsilon$ ).

### 14.1 Follow-Up: Why is the boundary layer at 0?

$$
\begin{aligned}
\epsilon y^{\prime \prime}+2 y^{\prime}+y & =0, \quad 0<x<1 \\
y(0) & =0 \\
y(1) & =1
\end{aligned}
$$

Try to find the solution with the boundary layer at $x=1$.
(a) Outer solution.

$$
\begin{aligned}
y & =y_{0}+\epsilon y_{1}(x)+\cdots \\
2 y_{0}^{\prime}+y_{0} & =0, \quad 0<x<1 \\
y_{0}(0) & =0 \\
y_{0} & =c e^{-x / 2} \quad \Rightarrow \quad y_{0}=0
\end{aligned}
$$

(b) Inner solution near $x=1$.

$$
\begin{aligned}
X & =\frac{1-x}{\epsilon} \\
y(x ; \epsilon) & =Y(X ; \epsilon) \\
\frac{d}{d x} & =-\frac{1}{\epsilon} \frac{d}{d X} \\
Y^{\prime \prime}-2 Y^{\prime}+\epsilon Y & =0, \quad 0<X<\infty \quad\left(Y^{\prime}=\frac{d Y}{d X}\right) \\
Y(0) & =1 \\
Y & =Y_{0}+\epsilon Y_{1}+\cdots \\
Y_{0}^{\prime \prime}-2 Y_{0}^{\prime} & =0 \\
Y_{0}(0) & =1 \\
Y_{0}(X) & =c_{1}+c_{2} e^{2 X} \\
& =1+c\left(1-e^{2 x}\right)
\end{aligned}
$$

(c) Matching. We want $y_{0}(x)$ as $x \rightarrow 1^{-}$to match with $Y_{0}(X)$ as $X \rightarrow \infty$.

$$
\begin{aligned}
& y_{0}(x) \rightarrow 0 \quad \text { as } x \rightarrow 1^{-} \\
& Y_{0}(x) \rightarrow\left\{\begin{array}{rr}
\infty & c>0 \\
1 & c=0 \\
-\infty & c<0
\end{array}\right.
\end{aligned}
$$

So after going through all of this analysis, we find that it won't work.

### 14.2 General Linear 2nd Order BVP's

$$
\begin{aligned}
\epsilon y^{\prime \prime}+a(x) y^{\prime}+b(x) y & =0, \quad 0<x<1 \\
y(0) & =\alpha \\
y(1) & =\beta
\end{aligned}
$$

Find an asymptotic solution as $\epsilon \rightarrow 0^{+}$. Suppose $a(x) \geq \delta>0$ on $0 \leq x \leq 1$.
Claim: we get a boundary layer at $x=0$.

1. $X=\frac{x}{\epsilon}$. The leading order inner equation for $Y_{0}$ is

$$
\begin{aligned}
Y_{0}^{\prime \prime}+a(0) Y_{0}^{\prime} & =0 \\
Y_{0}(X) & =c_{1}+c_{2} e^{-a(0) X} \\
& \rightarrow c_{1} \quad \text { as } X \rightarrow \infty \text { if } a(0)>0
\end{aligned}
$$

2. $X=\frac{1-x}{\epsilon}$ for a boundary layer at $x=1$.

$$
\begin{aligned}
Y_{0}^{\prime \prime}-a(1) Y_{0}^{\prime} & =0 \\
Y_{0}(X) & =c_{1}+c_{2} e^{a(1) X}
\end{aligned}
$$

We need $a(1)<0$ in order to permit matching.

So

1. If $a(x) \geq \delta>0$ we get a boundary layer at $x=0$.
2. If $a(x) \leq-\delta<0$ we get a boundary layer at $x=1$

If $a(x)$ changes sign (turning points), we get more complicated behavior.
3 . If $a(0)<0, a(1)>0$, we get no boundary layers (maybe interior/corner layer).
4. If $a(0)>0, a(1)<0$, we can have boundary layers at both endpoints.

### 14.2.1 Boundary Layer Example 1

$$
\begin{aligned}
& \epsilon y^{\prime \prime}+x y^{\prime}-y=0, \quad-1<x<1 \\
& y(-1)=1 \\
& y(1)=2 \\
&\left\{\begin{array}{l}
a(-1)=-1<0 \\
a(1)=1>0
\end{array} \Rightarrow\right. \text { no BL possible at either endpoint }
\end{aligned}
$$

(a) Outer solution.

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\cdots \\
x y_{0}^{\prime}-y_{0} & =0 \\
y_{0}(x) & =C x
\end{aligned}
$$

Impose left and right boundary conditions to get left and right outer solutions.

$$
\begin{aligned}
y_{0}^{L}(x) & =-x \\
y_{0}^{R}(x) & =2 x
\end{aligned}
$$

Try

$$
y_{0}(x)=\left\{\begin{array}{cl}
-x & -1 \leq x<0 \\
2 x & 0<x \leq 1
\end{array}\right.
$$


(b) Inner solution. Introduce scaled variable

$$
\begin{aligned}
X & =\frac{x}{\delta} \\
y(x) & =\delta Y(X) \\
\frac{d}{d x} & =\frac{1}{\delta} \frac{d}{d X} \\
x & =\delta X \\
\frac{\epsilon}{\delta^{2}} Y^{\prime \prime}+\delta X \cdot \frac{1}{\delta} Y^{\prime}-Y & =0 \\
\frac{\epsilon}{\delta^{2}} Y^{\prime \prime}+X Y^{\prime}-Y & =0
\end{aligned}
$$

We have a dominant three-term balance for $\delta=\epsilon^{1 / 2}$.

$$
Y^{\prime \prime}+X Y^{\prime}-Y=0, \quad-\infty<X<\infty
$$

Matching.

$$
\begin{aligned}
y_{0}^{L}(x) & =-\delta\left(\frac{x}{\delta}\right)=-\delta X \\
y_{0}^{R}(x) & =\delta\left(\frac{2 x}{\delta}\right)=\delta 2 X \\
Y(X) & \sim-X \quad \text { as } X \rightarrow-\infty \\
Y(x) & \sim 2 X \quad \text { as } X \rightarrow \infty
\end{aligned}
$$

## $15 \quad 5-4-12$

### 15.1 Boundary Layers (Continued)

$$
\begin{aligned}
\epsilon y^{\prime \prime}+a(x) y^{\prime}+b(x) y & =0 \\
y(0) & =\alpha \\
y(1) & =\beta
\end{aligned}
$$

A boundary layer at $x=0$ is possible if $a(0)>0$, and a boundary layer at $x=1$ is possible if $a(1)<0$. If $a(x)$ changes signs, more complications may occur.

### 15.1.1 Boundary Layer Example 1 (From Last Time)

$$
\begin{aligned}
\epsilon y^{\prime \prime}+x y^{\prime}-y & =0, \quad-1<x<1 \\
y(-1) & =1 \\
y(1) & =2
\end{aligned}
$$

There was no way to put in a boundary layer at either endpoint because as $x$ changes signs you change from growing to decaying solutions.

Outer solution:

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\cdots \\
x y_{0}^{\prime}-y_{0} & =0 \\
y_{0}(x) & =C x \\
y_{0}^{L}(x) & =-x \\
y_{0}^{R}(x) & =2 x
\end{aligned}
$$



The simplest, where we have a corner layer at $x=0$, is the right solution because it can be matched.
Inner solution: (for the corner layer)

$$
\begin{aligned}
y & =\epsilon^{1 / 2} Y\left(\frac{x}{\epsilon^{1 / 2}}\right) \\
X & =\frac{x}{\epsilon^{1 / 2}}
\end{aligned}
$$

Here we have a 3 -term dominant balance, and we get

$$
Y_{0}^{\prime \prime}+x Y_{0}^{\prime}-Y_{0}=0
$$

and then we have to subject this to the matching conditions.
Matching conditions:
inner limit of outer solution $=$ outer limit of inner solution

$$
\begin{aligned}
y_{0}^{L}(x) & =-x \\
& =-\epsilon^{1 / 2} \frac{x}{\epsilon^{1 / 2}} \\
& =-\epsilon^{1 / 2} X
\end{aligned}
$$

The solution

$$
Y_{0}(X)=c_{1} X+c_{2}\left[e^{-\frac{1}{2} X}+X \int_{-\infty}^{x} e^{-t^{2} / 2} d t\right]
$$

as $X \rightarrow-\infty$, and this looks like $c_{1} X$, so let $c_{1}=-1$. As $X \rightarrow \infty$,

$$
\begin{aligned}
Y_{0}(X) & \sim\left[c_{1}+c_{2} \int_{-\infty}^{\infty} e^{-t^{2} / 2} d t\right] x \\
c_{2} & =\frac{3}{\sqrt{2 \pi}}
\end{aligned}
$$

Question: what is the uniform solution? It would look like

$$
\begin{aligned}
& y \sim y_{\text {inner }}+y_{\text {outer }}^{L}+y_{\text {outer }}^{R}-y_{\text {overlap }}^{L}-y_{\text {overlap }}^{R} \\
& y \sim-x+\frac{3 \epsilon^{1 / 2}}{\sqrt{2 \pi}} e^{-x^{2} / 2 \epsilon}+\frac{3}{\sqrt{2 \pi}} x \int_{-\infty}^{x / \epsilon^{1 / 2}} e^{-t^{2} / 2} d t
\end{aligned}
$$

More important than using the inner solution is that it matches with respect to the boundaries and outer solution.

### 15.1.2 Boundary Layer Example 2

$$
\begin{aligned}
\epsilon y^{\prime \prime}-x y^{\prime}+y & =0, \quad-1<x<1 \\
y(-1) & =1 \\
y(1) & =2
\end{aligned}
$$

So here $a(x)=-x, a(-1)=1$, and $a(1)=-1$ (so boundary layers are possible at both $x=-1$ and $x=1$ ).
Outer solution: (away from any boundary layers)

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\cdots \\
-x y_{0}^{\prime}+y_{0} & =0 \\
y_{0}(x) & =c x
\end{aligned}
$$

We'll leave $c$ arbitrary since it is not clear which BC to impose.

Inner solution at $x=-1$ :

$$
\begin{aligned}
X & =\frac{x+1}{\epsilon} \\
y(x ; \epsilon) & =Y(X ; \epsilon) \\
\frac{d}{d x} & =\frac{1}{\epsilon} \frac{d}{d X} \\
x & =-1+\epsilon X \\
\frac{1}{\epsilon} Y^{\prime \prime}-(-1+\epsilon X) \frac{1}{\epsilon} Y^{\prime}+Y & =0 \\
Y(0 ; \epsilon) & =1 \\
Y & =Y_{0}(X)+\epsilon Y_{1}(X)+\cdots \\
Y_{0}^{\prime \prime}+Y_{0}^{\prime} & =0 \\
Y_{0}(0) & =1 \\
Y_{0}(X) & =1+A\left(1-e^{-X}\right)
\end{aligned}
$$

$\underline{\text { Matching condition at } x=1}$

$$
\begin{aligned}
\lim _{X \rightarrow \infty} Y(X) & =\lim _{x \rightarrow-1} y_{0}(x) \\
1+A & =-c
\end{aligned}
$$

### 16.1 Boundary Layer Example 2

$$
\begin{aligned}
\epsilon y^{\prime \prime}-x y^{\prime}+y & =0, \quad-1<x<1 \\
y(-1) & =1 \\
y(1) & =2
\end{aligned}
$$

Boundary layers are possible at both endpoints.
Outer expansion:

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\cdots \\
-x y_{0}^{\prime}+y_{0} & =0 \\
y_{0}(x) & =C x
\end{aligned}
$$

Inner expansion $(x=-1)$ :

$$
\begin{aligned}
X & =\frac{x+1}{\epsilon} \quad\left(=\frac{x-1}{\delta}\right) \\
Y(X ; \epsilon) & =y(x ; \epsilon) \\
Y & =Y_{0}(X)+\epsilon Y_{1}(X)+\cdots \\
Y_{0}^{\prime \prime}+Y_{0}^{\prime} & =0 \\
Y_{0}(X) & =1+A\left(1-e^{-X}\right) \quad\left(Y_{0}(0)=1\right)
\end{aligned}
$$

Matching at $x=-1$ :

$$
\begin{aligned}
\lim _{x \rightarrow-1^{+}} y_{0}(x) & =\lim _{X \rightarrow \infty} Y_{0}(X) \\
-C & =1+A
\end{aligned}
$$



Inner expansion $(x=1)$ :

$$
\begin{aligned}
X & =\frac{1-x}{\epsilon} \\
Y(X ; \epsilon) & =y(x ; \epsilon) \\
\frac{d}{d x} & =-\frac{1}{\epsilon} \frac{d}{d X} \\
\frac{1}{\epsilon} Y^{\prime \prime}+\frac{1}{\epsilon}(1+\epsilon X) Y^{\prime}+Y & =0, \quad Y(0 ; \epsilon)=2 \\
Y & =Y_{0}(X)+\epsilon Y_{1}(X)+\cdots \\
Y_{0}^{\prime \prime}+Y_{0}^{\prime} & =0, \quad Y_{0}(0)=2 \\
Y_{0}(X) & =2+B\left(1-e^{-X}\right)
\end{aligned}
$$

Matching:

$$
\begin{aligned}
& \lim _{x \rightarrow 1} y_{0}(x)=\lim _{X \rightarrow \infty} Y_{0}(X) \\
& C=2+B
\end{aligned}
$$

So the solution is

$$
\begin{aligned}
y & \sim\left\{\begin{aligned}
-1+A\left[1-e^{-(1+x) / \epsilon}\right] \\
C x \\
2+B\left[1-e^{-(1-x) / \epsilon}\right]
\end{aligned}\right. \\
-C & =1+A \\
C & =2+B
\end{aligned}
$$

The problem is that $C$ is undetermined. It remains undetermined to all orders in $\epsilon$.
We can find $C$ here by using symmetry of the problem.

$$
\begin{aligned}
y(x) & =\frac{1}{2} x+z(x) \\
\epsilon z^{\prime \prime}-x\left(\frac{1}{2}+z^{\prime}\right)+\frac{1}{2} x+z & =0 \\
\epsilon z^{\prime \prime}-x z^{\prime}+z & =0 \\
z(-1) & =\frac{3}{2} \\
z(1) & =\frac{3}{2}
\end{aligned}
$$

This is invariant under $x \rightarrow-x, z \rightarrow z$. So for a solution $y=\frac{1}{2} x+z$ (assuming it's unique), $z$ is an even function of $x$.

$$
\begin{aligned}
y & \sim\left\{\begin{aligned}
-C-A e^{-(1+x) / \epsilon} \\
C x \\
C-B e^{-(1-x) / \epsilon}
\end{aligned}\right. \\
-C & =1+A \\
C & =2+B \\
C & =\frac{1}{2} \\
A & =B=-\frac{3}{2}
\end{aligned}
$$

This holds in the leading order solution if $C=\frac{1}{2}$, which implies that $A=B=-\frac{3}{2}$.

$$
y(x) \sim\left\{\begin{aligned}
\frac{1}{2}+\frac{3}{2} e^{-(1+x) / \epsilon} & 1+x=O(\epsilon) \\
\frac{1}{2} x & -1<x<1 \\
\frac{1}{2}+\frac{3}{2} e^{-(1-x) / \epsilon} & 1-x=O(\epsilon)
\end{aligned}\right.
$$

The uniform solution would be

$$
\begin{aligned}
y_{\text {uniform }} & \sim-\frac{1}{2}+\frac{3}{2} e^{-(1+x) / \epsilon}+\frac{1}{2} x+\frac{1}{2}+\frac{3}{2} e^{-(1-x) / \epsilon}-\left(-\frac{1}{2}\right)-\frac{1}{2} \\
& =\frac{1}{2} x+\frac{3}{2}\left[e^{-(1+x) / \epsilon}+e^{-(1-x) / \epsilon}\right]
\end{aligned}
$$

### 16.2 Boundary Layer Example 3

$$
\begin{aligned}
\epsilon y^{\prime \prime}-y y^{\prime}+y & =0, \quad 0<x<1 \\
y(0) & =1 \\
y(1) & =-1
\end{aligned}
$$

A comparison with the linear equation suggests no boundary layer at $x=0$ or $x=1$.

## $17 \quad 5-9-12$

### 17.1 Boundary Layer Example 3

$$
\begin{aligned}
\epsilon y^{\prime \prime}-y y^{\prime}+y & =0, \quad 0<x<1 \\
y(0) & =1 \\
y(1) & =-1
\end{aligned}
$$

Look for a solution with no boundary layers at $x=0$ or $x=1$.
Outer solution:

$$
\begin{aligned}
y & =y_{0}(x)+\epsilon y_{1}(x)+\cdots \\
-y_{0} y_{0}^{\prime}+y_{0} & =0 \\
y_{0}\left(-y_{0}^{\prime}+1\right) & =0
\end{aligned}
$$

Either

$$
\begin{aligned}
& y_{0}=0 \\
& y_{0}^{\prime}=1, \quad y_{0} \quad=x+c
\end{aligned}
$$

The left outer solution is

$$
\begin{aligned}
y_{0}^{L}(x) & =x+1 \\
y_{0}^{L}(0) & =1
\end{aligned}
$$

The right outer solution is

$$
\begin{aligned}
y_{0}^{R}(x) & =x-2 \\
y_{0}^{R}(1) & =-1
\end{aligned}
$$



Look for an interior layer of width $O(\epsilon)$ where, at $x_{0}\left(0<x_{0}<1\right)$, the solution jumps from the left outer
solution to the right outer solution.

$$
\begin{array}{rlrl}
X & =\frac{x-x_{0}}{\epsilon} \\
Y(X ; \epsilon) & =y(x ; \epsilon) \\
\frac{d}{d x} & =\frac{1}{\epsilon} \frac{d}{d X} \\
Y^{\prime \prime}-Y Y^{\prime}+\epsilon Y & =0 \\
Y & =Y_{0}(X)+\epsilon Y_{1}(X)+\cdots \\
Y_{0}^{\prime \prime}-Y_{0} Y_{0}^{\prime} & =0 \\
Y_{0}^{\prime}-\frac{1}{2} Y_{0}^{2} & =k \\
Y_{0}^{\prime} & =k+\frac{1}{2} Y_{0}^{2} \\
> & & k>0 \\
\gg & & k=0 \\
\gg & &
\end{array}
$$

Matching:

$$
\begin{aligned}
k & =-\frac{1}{2} a^{2}<0 \quad(a>0) \\
Y_{0}^{\prime} & =-\frac{1}{2} a^{2}+\frac{1}{2} Y_{0}^{2} \\
Y_{0}(X) & \rightarrow a \quad \text { as } X \rightarrow-\infty \\
Y_{0}(X) & \rightarrow-a \quad \text { as } X \rightarrow \infty
\end{aligned}
$$

This requires that $x_{0}=\frac{1}{2}$ in order to jump from $-a$ to $a$.
Matching condition:

$$
\begin{aligned}
& \lim _{X \rightarrow \infty} Y_{0}(X)=\lim _{x \rightarrow x_{0}^{+}} y_{0}^{R}(x)-a=-\frac{3}{2} \\
& \lim _{X \rightarrow-\infty} Y_{0}(X)=\lim _{x \rightarrow x_{0}^{-}} y_{0}^{L}(x) \\
& a=\frac{3}{2}
\end{aligned}
$$

So $a=\frac{3}{2}$. The solution is

$$
Y_{0}(x)=-\frac{3}{2} \tanh \left[\frac{3}{4}(X-c)\right]
$$

This constant $c$ is left undetermined (to all orders in $\epsilon$ ). Note that the system is invariant under $x \rightarrow 1-x$, $y \rightarrow-y$ (and the boundary conditions also remain unchanged). So the solution (if unique) must be odd about $x=\frac{1}{2}$. So $y\left(\frac{1}{2}\right)=0$ and therefore $c=0$.

Summary:

$$
y \sim\left\{\begin{array}{rl}
x+1 & 0 \leq x<\frac{1}{2} \\
-\frac{3}{2} \tan \left[\frac{3\left(x-\frac{1}{2}\right)}{4 \epsilon}\right] & x-\frac{1}{2}=O(\epsilon) \\
x-2 & \frac{1}{2}<x \leq 1
\end{array}\right.
$$

The uniform (composite) solution is

$$
y(x) \sim x-\frac{1}{2}-\frac{3}{2} \tan \left[\frac{3\left(x-\frac{1}{2}\right)}{4 \epsilon}\right]
$$

## $18 \quad 5-11-12$

### 18.1 Method of Multiple Scales (MMS) and Oscillations



Pendulum

$$
\ddot{x}+\sin x=0
$$

Linearized equation at $x=0$ :

$$
\begin{aligned}
\ddot{x}+x & =0 \quad \text { (simple harmonic oscillator) } \\
x(t) & =a \cos t+b \sin t \\
& =A e^{i t}+A^{*} e^{-i t}, \\
A & =\frac{a-i b}{2}
\end{aligned}
$$

Look for small-amplitude solutions of the nonlinear equation (weakly nonlinear). Introduce a small parameter $\epsilon>0$ and look for solutions

$$
x(t, \epsilon)=\epsilon x_{1}(t)+\epsilon^{3} x_{2}(t)+\epsilon^{5} x_{3}(t)+O\left(\epsilon^{7}\right)
$$

For example, we could have

$$
\begin{aligned}
x(0, \epsilon) & =\epsilon \\
\dot{x}(0, \epsilon) & =0 \\
\sin x & =x-\frac{1}{6} x^{3}+O\left(x^{5}\right) \\
& =\epsilon x_{1}+\epsilon^{3} x_{2}-\frac{1}{6} \epsilon^{3} x_{1}^{3}+O\left(\epsilon^{5}\right) \\
\epsilon \ddot{x}_{1}+\epsilon^{3} \ddot{x}_{2}+\epsilon x_{1}+\epsilon^{3}\left(x_{2}-\frac{1}{6} x_{1}^{3}\right)+O\left(\epsilon^{5}\right) & =0 \\
O(\epsilon): \quad \ddot{x}_{1}+x_{1} & =0 \\
O\left(\epsilon^{3}\right): \quad \ddot{x}_{2}+x_{2} & =\frac{1}{6} x_{1}^{3} \\
x_{1}(t) & =A e^{i t}+A^{*} e^{-i t} \\
& =A e^{i t}+\underbrace{\text { c.c. }}_{\text {complex conjugate }} \\
\ddot{x}_{2}+x_{2} & =\frac{1}{6}\left[A e^{i t}+A^{*} e^{-i t}\right]^{3} \\
& =\frac{1}{6}\left[A^{3} e^{3 i t}+3|A|^{2} A e^{i t}+3|A|^{2} A^{*} e^{-i t}+\left(A^{*}\right)^{3} e^{-3 i t}\right]
\end{aligned}
$$

Side calculation: the solution of

$$
\begin{aligned}
\ddot{y}+y & =C e^{3 i t} \\
y(t) & =D e^{3 i t} \\
\ddot{y}+y & =(-9+1) D e^{3 i t} \\
& =-8 D e^{3 i t} \\
D & =-\frac{1}{8} C
\end{aligned}
$$

Another side calculation: consider

$$
\ddot{y}+y=C e^{i t}
$$

$e^{i t}$ is a solution of the homogeneous equation, so try

$$
\begin{aligned}
y(t) & =D t e^{i t} \\
\dot{y} & =D(i t+1) e^{i t} \\
\ddot{y} & =D(-t+i) e^{i t}+i D e^{i t} \\
& =D(-t+2 i) e^{i t} \\
\ddot{y}+y & =2 i D e^{i t} \\
D & =\frac{C}{2 i}
\end{aligned}
$$

Back to our problem, we have

$$
x_{2}(t)=-\frac{A^{3}}{48} e^{3 i t}+\frac{|A|^{2} A}{4 i} t e^{i t}-\frac{|A|^{2} A^{*}}{4 i} t e^{i t}-\frac{\left(A^{*}\right)^{3}}{48} e^{-3 i t}+B e^{i t}+B^{*} e^{-i t}
$$

Note: terms like $t e^{i t}$ appear in $x_{2}(t)$. The actual solution is a periodic function of time! Terms like $t e^{i t}$ are called secular terms.

The perturbation expansion becomes invalid when $t=O\left(1 / \epsilon^{2}\right)$ and $\epsilon^{2} x_{2}=O\left(\epsilon x_{1}\right)$.

### 18.1.1 Example

The origin of secular terms is the change in period/frequency of nonlinear oscillations with amplitude:

$$
\begin{aligned}
\epsilon \cos \left(\left(1+\epsilon^{2}\right) t\right) & =\epsilon \cos \left(t+\epsilon^{2} t\right) \\
& =\epsilon \cos t-(\sin t) \epsilon^{3} t+O\left(\epsilon^{4}\right)
\end{aligned}
$$

There is a nonuniformity in the expansion as $\epsilon \rightarrow 0$ for large $t$. In a sense, the largeness of $t$ overcomes the smallness of $\epsilon$.

### 18.2 Poincaré-Lindstedt Method

Introduce a rescaled time,

$$
\tau=\omega(\epsilon) t
$$

Expand the frequency as

$$
\omega(\epsilon)=1+\epsilon^{2} \omega_{2}+\cdots
$$

Choose $\omega_{2}$ to ensure that no secular terms appear.

## $19 \quad 5-14-12$

### 19.1 Poincaré-Lindstedt Method

Pendulum:

$$
\ddot{x}+\sin x=0
$$

We want to obtain an asymptotic solution for small-amplitude periodic solutions. Straightforward expansion fails due to secular terms (from dependence of the period on amplitude).

Idea: introduce a "strained" time

$$
\begin{aligned}
\tau & =\omega t \\
x(t) & =y(\omega t)=y(\tau)
\end{aligned}
$$

Recall that $y(\tau)$ is $2 \pi$-periodic in $\tau$. The $2 \pi$ is for convenience. The important point is that the period of $y(\tau)$ is fixed.

$$
\begin{aligned}
\frac{d}{d t} & =\omega \frac{d}{d \tau} \\
\dot{x} & =\omega \dot{y}, \quad \dot{y}=\frac{d y}{d \tau} \\
\omega^{2} \ddot{y}+\sin y & =0
\end{aligned}
$$

Expand:

$$
\begin{aligned}
y & =\epsilon y_{1}(\tau)+\epsilon^{3} y_{2}(\tau)+\cdots \\
\omega & =\omega_{0}+\epsilon^{2} \omega_{1}+\cdots \\
y(\tau+2 \pi) & =y(\tau) \\
\sin y & =y-\frac{1}{6} y^{3}+O\left(y^{5}\right) \\
& =\epsilon y_{1}+\epsilon^{3} y_{2}-\frac{1}{6} \epsilon^{3} y_{1}^{3}+O\left(\epsilon^{5}\right) \\
2 \epsilon^{2} \omega_{0} \omega_{1} & \leftarrow \epsilon \omega_{0}^{2} \ddot{y}_{1}+\epsilon^{3}\left[\omega_{0}^{2} \ddot{y}_{2}+2 \omega_{0} \omega_{1} \ddot{y}_{1}\right]+\cdots \\
\left(\omega_{0}^{2}+2 \epsilon^{2} \omega_{0} \omega_{1}+\cdots\right)\left(\epsilon \ddot{y}_{1}+\epsilon^{3} \ddot{y}_{2}+\cdots\right)+\epsilon y_{1}+\epsilon^{3}\left(y_{2}-\frac{1}{6} y_{1}^{3}\right) & =O\left(\epsilon^{5}\right) \\
O(\epsilon): \quad \omega_{0}^{2} \ddot{y}_{1}+y_{1} & =0 \\
y_{1}(\tau+2 \pi) & =y_{1}(\tau) \\
\omega_{0}^{2} \ddot{y}_{2}+y_{2} & =\frac{1}{6} y_{1}^{3}-2 \omega_{0} \omega_{1} \ddot{y}_{1} \\
y_{2}(\tau+2 \pi) & =y_{2}(\tau)
\end{aligned}
$$

From the leading order equation, we need $\omega_{0}^{2}=1\left(\omega_{0}=1\right)$. Then

$$
y_{1}(\tau)=A e^{i \tau}+A^{*} e^{-i \tau}
$$

Next order:

$$
\begin{aligned}
\ddot{y}_{2}+y_{2} & =\frac{1}{6} y_{1}^{3}-2 \omega_{1} \ddot{y}_{1} \\
y_{2}(\tau+2 \pi) & =y_{2}(\tau) \\
\ddot{y}_{2}+y_{2} & =\frac{1}{6}\left(A^{3} e^{3 i \tau}+3 A^{2} A^{*} e^{i \tau}+3 a\left(A^{*}\right)^{2} e^{-i \tau}+\left(A^{*}\right)^{3} e^{-3 i \tau}\right)+2 \omega_{1}\left(A e^{i \tau}+A^{*} e^{-i \tau}\right) \\
& =\frac{1}{6} A^{3} e^{3 i \tau}+\left[\frac{1}{2} A|A|^{2}+2 \omega_{1} A\right] e^{i t}+\left[\frac{1}{2} A^{*}|A|^{2}+2 \omega_{1} A^{*}\right] e^{-i \tau}+\frac{1}{6}\left(A^{*}\right)^{3} e^{-3 i \tau}
\end{aligned}
$$

The solution has the form

$$
y_{2}(\tau)=B e^{3 i \tau}+C \tau e^{i \tau}+\text { complex conjugates }
$$

$C \tau e^{i \tau}$ is a secular term (non-periodic), from the resonant term $\propto e^{i \tau}$ that solution of the homogeneous equation. We only get a periodic solution for $y_{2}(\tau)$ if the coefficient of $e^{i \tau}$ on the RHS is zero. So

$$
\begin{aligned}
\frac{1}{2}|A|^{2}+2 \omega_{1} & =0 \\
\omega_{1} & =-\frac{1}{4}|A|^{2} \\
\ddot{y}_{2}+y_{2} & =\frac{1}{6} A^{3} e^{3 i \tau}+\text { complex conjugates } \\
y_{2}(\tau) & =B e^{3 i \tau}+\text { complex conjugates } \\
-9 B+B & =\frac{1}{6} A^{3} \\
B & =-\frac{1}{48} A^{3} \\
y(\tau) & =A e^{i \tau}+\text { complex conjugate }-\frac{1}{48} \epsilon^{3} A^{3} e^{3 i \tau}+\text { complex conjugate }+O\left(\epsilon^{3}\right) \\
\omega & =1-\frac{1}{4} \epsilon^{2}|A|^{2}+O\left(\epsilon^{4}\right) \\
x(t ; \epsilon) & =y(\omega t ; \epsilon) \\
& =\epsilon A e^{i \omega t}-\frac{1}{48} \epsilon^{3} A^{3} e^{3 i \omega t}+\text { complex conjugate }+O\left(\epsilon^{5}\right) \\
\omega(\epsilon) & =1-\frac{1}{4} \epsilon^{2}|A|^{2}+O\left(\epsilon^{4}\right)
\end{aligned}
$$

For example, consider the solution with

$$
\left.\begin{array}{rl}
x=a \\
\dot{x}=0
\end{array}\right\} \quad \text { at } t=0, ~ \begin{aligned}
& =0 \\
\epsilon\left(A+A^{*}\right)-\frac{1}{48} \epsilon^{3}\left[A^{3}+\left(A^{*}\right)^{3}\right] & =a+\cdots \\
i \omega \epsilon\left(A-A^{*}\right)+\frac{1}{48} \cdot 3 i \omega \epsilon^{3}\left[A^{3}-\left(A^{*}\right)^{3}\right] & =0+\cdots \\
A & =A^{*} \text { is real } \\
2 \epsilon A-\frac{1}{24} \epsilon^{3} A^{3} & =a \\
\epsilon A & =\frac{1}{2} a+O\left(\epsilon^{3}\right) \\
& =\frac{1}{2} a+\frac{1}{384} a^{3}+O\left(a^{5}\right)
\end{aligned}
$$

So we are solving

$$
\begin{aligned}
\ddot{x}+\sin x & =0 \\
x(0) & =a \\
\dot{x}(0) & =0 \\
x(t) & =\frac{1}{2} a e^{i \omega t}+\frac{1}{2} a e^{-i \omega t}+\frac{1}{384} a^{3}\left(e^{i \omega t}+e^{-i \omega t}\right)-\frac{1}{384} a^{3}\left(e^{3 i \omega t}+e^{-3 i \omega t}\right)+O\left(a^{5}\right) \\
x(t) & =a \cos (\omega t)+\frac{1}{192} a^{3}[\cos (\omega t)-\cos (3 \omega t)]+O\left(a^{5}\right) \\
\omega & =1-\frac{1}{16} a^{2}+O\left(a^{4}\right)
\end{aligned}
$$

The period of the solution is

$$
\begin{aligned}
T & =\frac{2 \pi}{\omega}=2 \pi\left(\frac{1}{1-\frac{1}{16} a^{2}+\cdots}\right) \\
& =2 \pi\left(1+\frac{1}{16} a^{2}+O\left(a^{4}\right)\right)
\end{aligned}
$$

### 20.1 Poincaré-Lindstedt Method

$$
\ddot{x}+x=\epsilon F(t, x, \dot{x})
$$

Look for periodic solutions.

$$
\begin{aligned}
\tau & =\omega t \\
\omega^{2} \frac{d^{2} x}{d \tau^{2}}+x & =\epsilon F\left(t, x, \omega \frac{d x}{d \tau}\right) \\
x(\tau+2 \pi ; \epsilon) & =x(\tau ; \epsilon) \\
x(\tau ; \epsilon) & =x_{0}(\tau)+\epsilon x_{1}(\tau)+\cdots \\
\omega & =\omega_{0}+\epsilon \omega_{1}+\cdots \\
\omega_{0}^{2} \frac{d^{2} x_{0}}{d \tau^{2}}+x_{0} & =0 \\
\omega_{0} & =1 \quad \text { to get } 2 \pi-\text { periodic solutions } \\
x_{0} & =A e^{i \tau}+A^{*} e^{-i \tau} \\
\frac{d^{2} x_{n}}{d \tau^{2}}+x_{n} & =f_{n}, \quad f_{n} \text { depends on } x_{0}, \ldots, x_{n-1} \text { and } \omega_{1}, \ldots, \omega_{n-1}
\end{aligned}
$$

This has the form

$$
\begin{aligned}
L x_{n} & =f_{n} \\
L & =\frac{d^{2}}{d \tau^{2}}+1 \quad \text { acting on } 2 \pi \text {-periodic functions } x_{n} \in L^{2}(\mathbb{T})
\end{aligned}
$$

$L$ is a self-adjoint (Sturm-Liouville) operator with periodic BC's.

$$
\begin{aligned}
\langle f, g\rangle & =\int_{0}^{2 \pi} \overline{f(\tau)} g(\tau) d \tau \\
\langle f, L g\rangle & =\langle L f, g\rangle
\end{aligned}
$$

The eigenvalues are

$$
\begin{array}{ll} 
& L \phi=\lambda \phi \\
& \\
\lambda_{0}=1 & \phi_{0}=1 \\
\lambda_{n}=-n^{2}+1 & \phi_{n}=e^{i n t}, e^{-i n t}
\end{array}
$$

For $f \in L^{2}(\mathbb{T})$, when is $L u=f$ solvable? If $L \phi=0$,

$$
\begin{aligned}
\langle\phi, L u\rangle & =\langle\phi, f\rangle \\
\langle L \phi, u\rangle & =\langle\phi, f\rangle \\
\langle\phi, f\rangle & =0
\end{aligned}
$$

Fredholm alternative: $L u=f, L^{*}=L$ is solvable only if

$$
\langle\phi, f\rangle=0 \quad \forall \phi \text { such that } L \phi=0
$$

(The eigenfunction expansion shows it is sufficient also.)

For $L=\frac{d^{2}}{d \tau^{2}}+1$,

$$
\begin{aligned}
L \phi & =0 \\
\phi & =c_{1} e^{i \tau}+c_{2} e^{-i \tau}
\end{aligned}
$$

The solvability condition is

$$
\left\langle e^{i \tau}, f\right\rangle=\left\langle e^{-i \tau}, f\right\rangle=0
$$

which says that the Fourier coefficients $\hat{f}_{1}$ and $\hat{f}_{-1}$ vanish.

$$
\begin{aligned}
L x_{0} & =0 \\
x_{0} & =A e^{i \tau}+A^{*} e^{-i \tau} \\
L x_{n} & =f_{n}\left(x_{0}, \ldots, x_{n-1}, \omega_{1}, \ldots, \omega_{n-1}\right) \\
x_{n} & =x_{n}^{(p)}+A_{n} e^{i \tau}+A_{n}^{*} e^{-i \tau}
\end{aligned}
$$

Determine $\omega_{n-1}$ and (possibly) $\left|A_{n-1}\right|$ from the solvability conditions for $x_{n}$.

### 20.2 Weakly Damped Simple Harmonic Oscillator

$$
\ddot{x}+\epsilon \dot{x}+x=0, \quad 0<\epsilon \ll 1
$$

Straightforward expansion:

$$
\begin{aligned}
x & =x_{0}(t)+\epsilon x_{1}(t)+\cdots \\
\ddot{x}_{0}+x_{0} & =0 \\
x_{0} & =A e^{i t}+A^{*} e^{-i t} \\
\ddot{x}_{1}+x_{1} & =-\dot{x}_{0} \\
\ddot{x}_{1}+x_{1} & =-i A e^{i t}+i A^{*} e^{-i t}
\end{aligned}
$$

Get $t e^{-i t}$ terms in $x_{1}$ (secular). Here, introducing a variable $\tau=\omega t$ and looking for periodic solutions in $\tau$ doesn't help!


The solutions look like $e^{r t}$.

$$
\begin{aligned}
r^{2}+\epsilon r+1 & =0 \\
r & =-\frac{\epsilon \pm \sqrt{\epsilon^{2}-4}}{2} \\
& =-\frac{\epsilon}{2} \pm i \sqrt{1-\frac{\epsilon^{2}}{4}}
\end{aligned}
$$

Basic idea: we have two time-scales

1. The period of oscillations, $O(1) \Rightarrow t=t$
2. The time-scale of the damping, $O\left(\frac{1}{\epsilon}\right) \Rightarrow T=\epsilon t$

Introduce two "multiple-scale" variables simultaneously. Look for solutions of the form

$$
x=x(t, T ; \epsilon)
$$

and treat $t$ and $T$ as independent variables. (Evaluate $T=\epsilon t$ at the end.) This seems crazy because we have replaced an ODE with a PDE.

## $21 \quad 5-18-12$

### 21.1 Weakly Damped Oscillator

$$
(\mathrm{ODE}) \quad \ddot{x}+\epsilon \dot{x}+x=0
$$

We want to obtain an asymptotic solution that is valid for long times, $t=O\left(\frac{1}{\epsilon}\right)$. Straightforward expansion for $x(t ; \epsilon)$ leads to secular terms. For the method of multiple scales, we will introduce two time variables: $t, T=\epsilon t$. Look for a solution of the form

$$
x(t ; \epsilon)=y(t, \epsilon t ; \epsilon) .
$$

Then

$$
\begin{aligned}
\dot{x}(t ; \epsilon) & =y_{t}(t, \epsilon t ; \epsilon)+\epsilon y_{T}(t, \epsilon t ; \epsilon) \\
\ddot{x} & =y_{t t}+2 \epsilon y_{t T}+\epsilon^{2} y_{T T} \\
\frac{d}{d t} & \rightarrow \frac{\partial}{\partial t}+\epsilon \frac{\partial}{\partial T} \quad \text { (derivative expansion) } \\
(\mathrm{PDE}) \quad y_{t t}+2 \epsilon y_{t T}+\epsilon^{2} y_{T T}+\epsilon\left(y_{t}+\epsilon y_{T}\right)+y & =0
\end{aligned}
$$

$x(t ; \epsilon)$ satisfies the ODE if and only if $y(t, T ; \epsilon)$ satisfies the PDE on $T=\epsilon t$. The idea of the method of multiple scales is to require that $y(t, T ; \epsilon)$ satisfies the PDE for all $(t, T)$. So we start by introducing a lot of freedom, requiring that $x(t ; \epsilon)=y(t, \epsilon t ; \epsilon)$, and then we take it away by saying that it must satisfy the PDE for all $(t, T)$.

Expand:

$$
\begin{aligned}
y(t, T ; \epsilon) & =y_{0}(t, T)+\epsilon y_{1}(t, T)+O\left(\epsilon^{2}\right) \\
O(1): y_{0, t t}+y_{0} & =0 \\
O(\epsilon): \quad y_{1, t t}+y_{1}+2 y_{0, t T}+y_{0, t} & =0 \\
y_{0}(t, T) & =A(T) e^{i t}+A^{*}(T) e^{-i t}
\end{aligned}
$$

$y_{1, t t}+y_{1}+2 i A_{T} e^{i t}+$ complex conjugate $+i A e^{i t}+$ complex conjugate $=0$

$$
\begin{aligned}
y_{1, t t}+y_{1}+i\left(2 A_{T}+A\right) e^{i t}-i\left(2 A_{T}^{*}+A^{*}\right) e^{-i t} & =0 \\
y_{1}(t, T) & =C t e^{i t} \\
y_{1, t t}+y_{1} & =C\left(-t e^{i t}+2 e^{i t}\right)+C t e^{i t}=2 i C e^{i t} \\
C & =-\left(A_{T}+\frac{1}{2} A\right) \\
y_{1}(t, T)= & -\left(A_{T}+\frac{1}{2} A\right) t e^{i t}+\text { complex conjugate } \\
& +B e^{i t}+\text { complex conjugate }
\end{aligned}
$$

We require that the $y_{n}(t, T)$ don't grow too fast in $t$ (e.g. bounded functions of $t$ or sublinear). We get that
$y_{1}(t, T)$ is a bounded (periodic) function of $t$ only if the coefficient of $e^{i t}$ vanishes:

$$
\begin{aligned}
2 A_{T}+A & =0 \\
A(T) & =A_{0} e^{-T / 2} \\
y_{0}(t, T) & =A_{0} e^{-T / 2} e^{i t}+A_{0}^{*} e^{-T / 2} e^{-i t} \\
x(t ; \epsilon) & =A_{0} e^{-\epsilon t / 2} e^{i t}+\text { complex conjugate }+O(\epsilon) \quad \text { for } t=O\left(\frac{1}{\epsilon}\right) \\
r^{2}+\epsilon r+1 & =0 \\
r & =-\frac{\epsilon}{2} \pm i \sqrt{1-\frac{1}{4} \epsilon^{2}}
\end{aligned}
$$

## 21.2 van der Pol Oscillator

We already looked at strong damping:

$$
\epsilon \ddot{x}+\left(x^{2}-1\right) \dot{x}+x=0 .
$$

Weak damping:

$$
\ddot{x}+\epsilon\left(x^{2}-1\right) \dot{x}+x=0 .
$$

Strong damping:

$$
\begin{aligned}
\dot{x} & =y \\
\epsilon \dot{y} & =x-\left(x^{2}-1\right) y \\
\text { Slow manifold: } \quad y & =\frac{x}{1-x^{2}}
\end{aligned}
$$



Figure 8: There is a limit cycle in here somewhere. This is why we use the Lienard variables... (See Figure 6.)
Weak damping:


Figure 9: We spiral into the limit cycle from the outside, and we spiral away from the limit cycle on the inside.

## 22.1 van der Pol Equation

$$
\ddot{x}+\epsilon\left(x^{2}-1\right) \dot{x}+x=0 \quad \text { (weak damping) }
$$

Multiple scale variables: $t, T=\epsilon t$. Look for a solution of the form

$$
\begin{aligned}
x(t ; \epsilon) & =y(t, \epsilon t ; \epsilon) \\
\frac{d}{d t} & \left.\rightarrow \frac{\partial}{\partial t}\right|_{T}+\left.\epsilon \frac{\partial}{\partial T}\right|_{t} \\
y_{t t}+2 \epsilon y_{t T}+\epsilon^{2} y_{T T}+\epsilon\left(y^{2}-1\right)\left(y_{t}+\epsilon y_{T}\right)+y & =0 \\
y_{t t}+\epsilon\left[2 y_{t T}+\left(y^{2}-1\right) y_{t}\right]+\epsilon^{2}\left[y_{T T}+\left(y^{2}-1\right) y_{T}\right]+y & =0 \\
y & =y_{0}(t, T)+\epsilon y_{1}(t, T)+O\left(\epsilon^{2}\right) \\
y_{0, t t}+y_{0} & =0 \\
y_{1, t t}+y_{1}+2 y_{0, t T}+\left(y_{0}^{2}-1\right) y_{0, t} & =0 \\
y_{0}(t, T) & =A(T) e^{i t}+A^{*}(T) e^{-i t} \\
y_{1, t t}+y_{1}+2\left[i A_{T} e^{i t}-i A_{T}^{*} e^{-i t}\right]+\left[A^{2} e^{2 i t}+2|A|^{2}+\left(A^{*}\right)^{2} e^{-2 i t}-1\right]\left[i A e^{i t}-i A^{*} e^{-i t}\right] & =0 \\
y_{1, t t}+y_{1}+i A^{3} e^{3 i t}+\left[2 i A_{T}+i|A|^{2} A-i A\right] e^{i t}+\operatorname{complex} \text { conjugate } & =0
\end{aligned}
$$

We require that $y_{1}(t, T)$ is a periodic function of "fast" time $t$. So we must have

$$
\begin{aligned}
A_{T}+\frac{1}{2}\left(|A|^{2}-a\right) A & =0 \\
A(T) & =r(T) e^{i \phi(T)} \\
A_{T} & =\left[r_{T}+i r \phi_{T}\right] e^{i \phi} \\
r_{T}+i r \phi_{T}+\frac{1}{2}\left(r^{2}-1\right) r & =0 \\
r_{T}+\frac{1}{2} r\left(r^{2}-1\right) & =0 \\
\phi_{T} & =0 \\
\phi & =\phi_{0}
\end{aligned}
$$

$$
\begin{aligned}
x(t ; \epsilon) & =A(\epsilon t) e^{i t}+\text { complex conjugate }+O(\epsilon) \\
& =r(\epsilon t) e^{i\left(t+\phi_{0}\right)}+\text { complex conjugate }+O(\epsilon) \quad \text { for times } t=O\left(\frac{1}{\epsilon}\right) \\
r & =0 \quad \Rightarrow \quad x=0 \quad \text { (equilibrium) } \\
r & =1 \quad \Rightarrow \quad x=2 \cos \left(t+\phi_{0}\right)
\end{aligned}
$$



Let's try to formulate an energy argument for this system. Energy equation:

$$
\begin{aligned}
\dot{x} \ddot{x}+\dot{x} x+\epsilon\left(x^{2}-1\right) \dot{x}^{2} & =0 \\
\frac{d}{d t}\left(\frac{1}{2} \dot{x}^{2}+\frac{1}{2} x^{2}\right) & =-\epsilon\left(x^{2}-1\right) \dot{x}^{2} \begin{cases}>0 & |x|<1 \text { (negative damping) } \\
<0 & |x|>1 \text { (positive damping) }\end{cases}
\end{aligned}
$$

For a periodic solution,

$$
\oint\left(x^{2}-1\right) \dot{x}^{2} d t=0
$$

For weak damping:

$$
\begin{aligned}
x(t) & =a \cos t \\
\int_{0}^{2 \pi}\left(a^{2} \cos ^{2} t-1\right) \cdot a^{2} \sin ^{2} t d t & =0 \\
\frac{a^{2}}{2 \pi} \int_{0}^{2 \pi} \cos ^{2} t \cdot \sin ^{2} t d t & =\frac{1}{2 \pi} \int_{0}^{2 \pi} \sin ^{2} t d t \\
\frac{1}{2 \pi} \int_{0}^{2 \pi} \sin ^{2} t d t & =\frac{1}{2} \\
\frac{1}{2 \pi} \int_{0}^{2 \pi}\left(\cos ^{2} t \sin ^{2} t\right) d t & =\frac{1}{2 \pi} \int_{0}^{2 \pi}\left(\sin ^{2} t-\sin ^{4} t\right) d t \\
& =\frac{1}{2}-\frac{3}{8} \\
& =\frac{1}{8} \\
\frac{a^{2}}{8} & =\frac{1}{2} \\
a & =2
\end{aligned}
$$

## $23 \quad 5-23-12$

### 23.1 Method of Averaging

$$
\begin{aligned}
x_{t} & =\epsilon f(x, t) \\
x(0) & =c \\
x & =\left(\begin{array}{c}
x_{1} \\
x_{2} \\
\vdots \\
x_{n}
\end{array}\right) \\
f(x, t+2 \pi) & =f(x, t)
\end{aligned}
$$

$f: \mathbb{R}^{n} \times \mathbb{R} \rightarrow \mathbb{R}^{n}, f$ is periodic in time.
We introduce multiple scale variables $t, T=\epsilon t$. Then

$$
\begin{aligned}
x(t ; \epsilon) & =\left.y(t, T ; \epsilon)\right|_{T=\epsilon t} \\
\frac{d}{d t} & \rightarrow \frac{\partial}{\partial t}+\epsilon \frac{\partial}{\partial T} \\
y_{t}+\epsilon y_{T} & =\epsilon f(y, t)
\end{aligned}
$$

We look for solutions that are periodic in $t$ (i.e. no secular terms):

$$
\begin{aligned}
y(t+2 \pi, T ; \epsilon) & =y(t, T ; \epsilon) \\
y & =y_{0}(t, T)+\epsilon y_{1}(t, T)+O\left(\epsilon^{2}\right) \\
y_{0, t}+\epsilon y_{1, t}+\epsilon y_{0, T} & =\epsilon f\left(y_{0}, t\right)+O\left(\epsilon^{2}\right) \\
O(1): \quad y_{0, t} & =0 \\
y_{0} & =y_{0}(T) \\
O(\epsilon): \quad y_{1, t}+y_{0, T} & =f\left(y_{0}, t\right) \\
y_{1}(t+2 \pi, T) & =y_{1}(t, T) \\
0=\int_{0}^{2 \pi} y_{t} d t & =\int_{0}^{2 \pi} g(t) d t \\
\text { Need: } \bar{g} & =\frac{1}{2 \pi} \int_{0}^{2 \pi} g(t) d t=0
\end{aligned}
$$

We have

$$
y_{1, t}=-y_{0, T}+f\left(y_{0}, t\right)
$$

The solvability condition is that

$$
\begin{aligned}
\frac{1}{2 \pi} \int_{0}^{2 \pi}\left(-y_{0, T}+f\left(y_{0}, t\right)\right) d t & =0 \\
y_{0, T} & =\bar{f}\left(y_{0}\right) \\
\bar{f}\left(y_{0}\right) & =\frac{1}{2 \pi} \int_{0}^{2 \pi} f\left(y_{0}, t\right) d t \\
y(t) & =y_{0}(\epsilon t) \\
\partial_{T} & =\frac{1}{\epsilon} \partial_{t} \\
y_{t} & =\epsilon \bar{f}(y) \\
\bar{f}(y) & =\frac{1}{2 \pi} \int_{0}^{2 \pi} f(y, t) d t \\
x_{t} & =\epsilon f(x, t)
\end{aligned}
$$



Theorem 23.1.

For smooth $t$-periodic vector fields $f(x, t)$ there exist constants $\epsilon_{0}, c, k>0$ such that for all $\epsilon$ with $|\epsilon|<\epsilon_{0}$ we have

$$
|x(t ; \epsilon)-y(t)|<k \epsilon
$$

for $|t|<\frac{c}{\epsilon}$.

### 23.2 Geometrical Interpretation

$$
\begin{aligned}
x_{t} & =\epsilon f(x, t) \\
p^{\epsilon} & : \mathbb{R}^{n} \rightarrow \mathbb{R}^{n} \quad \text { (Poincaré map) } \\
x(0) & \mapsto x(2 \pi) \\
p^{\epsilon}\left(x_{0}\right)-x_{0} & =O(\epsilon)
\end{aligned}
$$



Figure 10: Poincaré map.
The flow of the averaged equation approximates the Poincaré map of the full equation (on times $t=O\left(\frac{1}{\epsilon}\right)$ ). Hyperbolic fixed points of the averaged equation correspond to $2 \pi$-periodic solutions of the full equation (for $\epsilon$ sufficiently small) with the same stability.

### 23.3 Periodic Standard Form

$$
\begin{align*}
\ddot{y}+y & =\epsilon g(y, \dot{y}, t) \quad(2 \pi \text {-periodic }) \\
y(t) & =x_{1}(t) \cos t+x_{2}(t) \sin t \\
\dot{y}(t) & =-x_{1}(t) \sin t+x_{2}(t) \cos t  \tag{23.1}\\
\binom{y}{\dot{y}} & =\left(\begin{array}{cc}
\cos t & \sin t \\
-\sin t & \cos t
\end{array}\right)\binom{x_{1}}{x_{2}} \\
\ddot{y} & =-x_{1} \cos t-x_{2} \sin t-\dot{x}_{1} \sin t+\dot{x}_{2} \cos t \\
& =-y-\dot{x}_{1} \sin t+\dot{x}_{2} \cos t \\
-\dot{x}_{1} \sin t+\dot{x}_{2} \cos t & =\epsilon g\left(x_{1} \cos t+x_{2} \sin t,-x_{1} \sin t+x_{2} \cos t, t\right)=\epsilon f(x, t) \\
\dot{x}_{1} \cos t+\dot{x}_{2} \sin t & =0 \quad(\text { so } 23.1 \text { holds }) \\
\dot{x}_{1} & =-\epsilon(\sin t) f(x, t) \\
\dot{x}_{2} & =\epsilon(\cos t) f(x, t) \\
\dot{x} & =\epsilon f(x, t)
\end{align*}
$$

## $24 \quad 5-25-12$

### 24.1 WKB Method

Simple harmonic oscillator with slowly varying frequency:

$$
x_{t t}+\omega^{2}(\epsilon t) x=0
$$



Figure 11: A pendulum system where the length of the pendulum can change.

$$
\begin{aligned}
T & =\epsilon t \\
\frac{d}{d t} & =\epsilon \frac{d}{d T} \\
\epsilon^{2} x_{T T}+\omega^{2}(T) X & =0
\end{aligned}
$$

Slow vs. small variations in frequency. Here, we use the fact that the variations are slow.
We want to find an approximate solution that is valid for $t=O\left(\frac{1}{\epsilon}\right)$. Try a multiple scale expansion: $t, T=\epsilon t$.

$$
\begin{aligned}
x(t ; \epsilon) & =\left.y(t, T ; \epsilon)\right|_{T=\epsilon t} \\
\frac{d}{d t} & \rightarrow \frac{\partial}{\partial t}+\epsilon \frac{\partial}{\partial T} \\
\frac{d^{2}}{d t^{2}} & \rightarrow \frac{\partial^{2}}{\partial t^{2}}+2 \epsilon \frac{\partial^{2}}{\partial t \partial T}+\epsilon^{2} \frac{\partial^{2}}{\partial T^{2}} \\
y_{t t}+2 \epsilon y_{t T}+\epsilon^{2} y_{T T}+\omega^{2}(T) y & =0 \\
y & =y_{0}(t, T)+\epsilon y_{1}(t, T)+\cdots \\
y_{0, t t}+\omega^{2}(T) y_{0} & =0 \\
y_{1, t t}+\omega^{2}(T) y_{1}+2 y_{0, t T} & =0 \\
y_{0}(t, T) & =A(T) e^{i \omega(T) t}+A^{*}(T) e^{-i \omega(T) t} \\
y_{0, t} & =i \omega A e^{i \omega t}+\operatorname{complex} \text { conjugate } \\
y_{0, t T} & =i(\omega A)_{T} e^{i \omega t}-\omega \omega_{T} A t e^{i \omega t}+\text { complex conjugate } \\
y_{1, t t}+\omega^{2} y_{1} & =2 \omega \omega_{T} A t e^{i \omega t}-i(\omega A)_{T} e^{i \omega t}+\text { complex conjugate }
\end{aligned}
$$

We get secular terms, and the solutions is not valid for long times $t=O\left(\frac{1}{\epsilon}\right)$.

Problem: the period is changing on a slow time-scale.
We've got oscillations with phase $\omega(T) t=\omega(\epsilon t) t$. The right way to do this is to use a "fast" phase

$$
\begin{aligned}
\theta & =\frac{\phi(\epsilon t)}{\epsilon} \\
\phi_{T}(T) & =\omega(T) .
\end{aligned}
$$

WKB expansion:

$$
\begin{aligned}
x(t ; \epsilon) & =\left.y(\theta, T ; \epsilon)\right|_{\theta=\frac{\phi(\epsilon t)}{\epsilon}, T=\epsilon t} \\
x(t ; \epsilon) & =y\left(\frac{\phi(\epsilon t)}{\epsilon}, \epsilon t ; \epsilon\right) \\
\frac{d x}{d t} & =\phi_{T} \frac{\partial y}{\partial \theta}+\epsilon \frac{\partial y}{\partial T} \\
\frac{d^{2} x}{d t^{2}} & =\phi_{T}\left[\phi_{T} \frac{\partial^{2} y}{\partial \theta^{2}}+\epsilon \frac{\partial^{2} y}{\partial T \partial \theta}\right]+\epsilon \phi_{T T} \frac{\partial y}{\partial \theta}+\epsilon\left[\phi_{T} \frac{\partial^{2} y}{\partial \theta \partial T}+\epsilon \frac{\partial^{2} y}{\partial T^{2}}\right] \\
& =\phi_{T}^{2} y_{\theta \theta}+\epsilon\left[\phi_{T T} y_{\theta}+2 \phi_{T} y_{\theta T}\right]+\epsilon^{2} y_{T T} \\
\phi_{T}^{2} y_{\theta \theta}+\epsilon\left[\phi_{T T} y_{\theta}+2 \phi_{T} y_{\theta T}\right]+\epsilon^{2} y_{T T}+\omega^{2}(T) y & =0
\end{aligned}
$$

Expand:

$$
y=y_{0}(\theta, T)+\epsilon y_{1}(\theta, T)+\cdots
$$

Require: $y(\theta, T ; \epsilon)$ is a $2 \pi$-periodic function of $\theta$.

$$
\begin{aligned}
\phi_{T}^{2} y_{0, \theta \theta}+\omega^{2} y_{0} & =0 \\
\phi_{T}^{2} y_{1, \theta \theta}+\omega^{2} y_{1}+\phi_{T T} y_{0, \theta}+2 \phi_{T} y_{0, \theta T} & =0
\end{aligned}
$$

$y_{0}(\theta, T)$ is $2 \pi$-periodic in $\theta$ if and only if $\phi_{T}^{2}=\omega^{2}$, or $\phi_{T}= \pm \omega$.

$$
\begin{aligned}
y_{0} & =A(T) e^{i \phi}+A^{*}(T) e^{-i \theta} \\
\omega^{2}\left(y_{1, \theta \theta}+y_{1}\right)+\phi_{T T}\left(i A e^{i \phi}+\text { c.c. }\right)+2 \phi_{T}\left(i A_{T} e^{i \phi}+\text { c.c. }\right) & =0 \\
\omega^{2}\left(y_{1, \theta \theta}+y_{1}\right)+\underbrace{i\left(2 \phi_{T} A_{T}+\phi_{T T} A\right)}_{=0 \text { so } y \text { is } 2 \pi \text {-periodic }} e^{i \phi}+\text { c.c. } & =0 \\
2 \phi_{T} A_{T}+\phi_{T T} A & =0 \\
\phi_{T} & =\omega \\
\left(\omega|A|^{2}\right)_{T} & =0
\end{aligned}
$$

## $25 \quad 5-30-12$

### 25.1 WKB Method

$$
\ddot{x}+\omega^{2}(\epsilon t) x=0
$$



$$
\begin{aligned}
& x(t ; \epsilon)=A(\epsilon t) e^{i \phi(\epsilon t) / \epsilon} \\
& T=\epsilon t \\
& \theta=\frac{\phi(\epsilon t)}{\epsilon} \\
& \dot{x}=\left(i \phi^{\prime} A+\epsilon A^{\prime}\right) e^{i \phi / \epsilon} \\
& \text { primes denote } \frac{d}{d T} \\
& \ddot{x}=i \phi^{\prime}\left(i \phi^{\prime} A+\epsilon A^{\prime}\right) e^{i \phi / \epsilon}+\left(\epsilon i \phi^{\prime \prime} A+\epsilon i \phi^{\prime} A^{\prime}+\epsilon^{2} A^{\prime \prime}\right) e^{i \phi / \epsilon} \\
&=\left[-\left(\phi^{\prime}\right)^{2} A+i \epsilon\left(2 \phi^{\prime}+\phi^{\prime \prime} A\right)+\epsilon^{2} A^{\prime \prime}\right] e^{i \phi / \epsilon} \\
& 0=-\left(\phi^{\prime}\right)^{2} A+i \epsilon\left(2 \phi^{\prime} A^{\prime}+\phi^{\prime \prime} A\right)+\epsilon^{2} A^{\prime \prime}+\omega^{2} A
\end{aligned}
$$

Choose $\left(\phi^{\prime}\right)^{2}=\omega^{2}$ to eliminate leading-order terms.

$$
2 \phi^{\prime} A^{\prime}+\phi^{\prime \prime} A=i \epsilon A^{\prime \prime}
$$

(Liouville-Green)
So far we haven't made any approximations. Let's look for an expansion

$$
\begin{aligned}
A & =A_{0}+\epsilon A_{1}+\epsilon^{2} A_{2}+\cdots \\
2 \phi^{\prime} A_{0}^{\prime}+\phi^{\prime \prime} A_{0} & =0
\end{aligned}
$$

Let's say we choose $\phi^{\prime}=\omega$.

$$
\begin{aligned}
A_{0}(T) & =\frac{1}{2} a(T) e^{i \delta} \\
2 \omega a^{\prime}+\omega^{\prime} a & =0 \\
\frac{a^{\prime}}{a} & =-\frac{\omega^{\prime}}{2 \omega} \\
\log a & =-\frac{1}{2} \log (\omega)+c \\
a & =\frac{a_{0}}{\sqrt{\omega}} \\
x & =A_{0}(T) e^{i \phi / \epsilon}+\text { complex conjugate }+O(\epsilon) \\
& =\frac{1}{2} a e^{i \delta} e^{i \phi / \epsilon}+\text { complex conjugate }+O(\epsilon) \\
x & =a \cos \left(\frac{\phi}{\epsilon}+\delta\right)+O(\epsilon) \\
\phi(T) & =\int_{0}^{T} \omega(\hat{T}) d \hat{T} \\
\omega a^{2} & =\operatorname{constant}
\end{aligned}
$$



$$
\begin{aligned}
x & =a(\epsilon t) \cos \left[\frac{\phi(\epsilon t)}{\epsilon}\right] \\
& =a\left(\epsilon t_{0}+\epsilon s\right) \cos \left[\frac{\phi\left(\epsilon t_{0}+\epsilon s\right)}{\epsilon}\right] \\
& =a\left(\epsilon t_{0}\right) \cos \left[\frac{1}{\epsilon}\left[\phi\left(\epsilon t_{0}\right)+\epsilon \phi^{\prime}\left(\epsilon t_{0}\right) s+O\left(\epsilon^{2}\right)\right]\right] \\
& \sim a\left(\epsilon t_{0}\right) \cos \left[\frac{\phi\left(\epsilon t_{0}\right)}{\epsilon}+\omega\left(\epsilon t_{0}\right) s\right]
\end{aligned}
$$

$$
t_{0}=O\left(\frac{1}{\epsilon}\right)
$$

$$
s=O(1)
$$

$$
t=t_{0}+s
$$

$\omega a^{2}$ is conserved under slow variations in $\omega$. For this reason, we say that $\omega a^{2}$ is adiabatic invariant, and we
call it the action.

$$
\begin{aligned}
\text { Energy } & =\frac{1}{2} \dot{x}^{2}+\frac{1}{2} \omega^{2} x^{2}=E \\
& =-a \phi^{\prime} \sin \left(\frac{\phi}{\epsilon}+\delta\right)+O(\epsilon) \\
& =-a \omega \sin \left(\frac{\phi}{\epsilon}+\delta\right)+O(\epsilon) \\
x & =a \cos \left(\frac{\phi}{\epsilon}+\delta\right)+O(\epsilon) \\
\text { Energy } & =\frac{1}{2} a^{2} \omega^{2}+O(\epsilon) \\
\text { Action } & =\frac{1}{2} \omega a^{2}=\frac{E}{\omega}
\end{aligned}
$$

There's an interesting quantum mechanical interpretation of the action involving energy levels.

### 25.2 Schrödinger Equation

$$
\begin{aligned}
i \hbar \Psi_{t} & =-\frac{\hbar}{2 m} \Psi_{x x}+V(x) \Psi \\
\Psi(x, t) & =\phi(x) e^{-i E t / \hbar} \\
-\frac{\hbar^{2}}{2 m} \phi_{x x}+V(x) \phi & =E \phi \\
\frac{\hbar^{2}}{2 m} \phi_{x x}+[E-V(x)] \phi & =0
\end{aligned}
$$

$\hbar \rightarrow 0$ corresponds to the WKB approximation, and this is called the semiclassical limit.

### 26.1 WKB Method and Turning Points

$$
\begin{aligned}
\epsilon^{2} y^{\prime \prime}+q(x) y & =0 \\
y & \sim a(x) e^{\phi(x) / \epsilon} \\
\left(\phi^{\prime}\right)^{2}+q & =0 \\
\phi^{\prime} & = \pm \sqrt{-q} \\
q>0 & \Rightarrow \quad \phi^{\prime}= \pm i \sqrt{q}, \quad \phi= \pm i S \\
y & \sim a e^{ \pm i S(x) / \epsilon} \\
q<0 & \Rightarrow \phi^{\prime}= \pm \sqrt{-q}, \quad \phi= \pm S \\
y & \sim a e^{ \pm S(x) / \epsilon}
\end{aligned}
$$

A turning point is where $q(x)=0, x \in \mathbb{R}$. At a simple zero ( $x=0$ is a turning point):

$$
q(x)=c x+O\left(x^{2}\right)
$$

the behavior changes from oscillatory to exponential. Airy equation:

$$
y^{\prime \prime}+x y=0
$$

The solutions are Airy functions: $A i(x)$ and $B i(x)$. Note: the $A$ stands for area, and $B$ follows $A$.
Let's say

$$
\begin{array}{rrr}
q(x)>0 & \text { when } x<x_{0} \\
q(x)<0 & \text { when } x>x_{0} \\
\epsilon^{2} y^{\prime \prime}+q(x) y=0 &
\end{array}
$$



Schrödinger equation:

$$
\begin{aligned}
i \hbar \Psi_{t} & =-\frac{\hbar^{2}}{2 m} \Psi_{x x}+V(x) \Psi \\
\Psi(x, t) & =\phi(x) e^{-i E t / \hbar} \\
-\frac{1}{2 m} \phi^{\prime \prime}+V(x) \phi & =E \phi \\
\phi^{\prime \prime}+2 m[E-V(x)] \phi & =0 \\
\phi(x) & =2 m[E-V(x)]
\end{aligned}
$$



### 26.2 A Model Bifurcation Problem for PDEs

$u(x, t)$ satisfies the following:

$$
\begin{aligned}
u_{t} & =u_{x x}+\mu \sin u, \quad 0<x<1, \quad t>0 \\
u(0, t) & =0 \\
u(1, t) & =0 \\
u(x, 0) & =f(x)
\end{aligned}
$$

This is a heat equation with a nonlinear heat source, $\mu \sin u . \mu \geq 0$ is a (dimensionless) parameter that measures the strength of the nonlinear heat sources.

Consider the equilibrium solution $u=0$. Is it stable?

1. We start by linearizing the PDE around $u=0$.

$$
\begin{aligned}
u_{t} & =u_{x x}+\mu u, \quad 0<x<1 \\
u(0, t) & =u(1, t)=0
\end{aligned}
$$

Separate variables.

$$
\begin{aligned}
u(x, t) & =e^{\sigma_{n} t} \sin (n \pi x), \quad n=1,2,3, \ldots \\
\sigma_{n} & =-n^{2} \pi^{2}+\mu
\end{aligned}
$$

$\sigma_{n}<0$ for all $n$ if $\mu<\pi^{2}$ ( $u=0$ is linearly stable). $\sigma_{1}>0$ if $\mu>\pi^{2}$ ( $u=0$ is linearly unstable).
2. How does the nonlinearity affect instability?

Assume $\mu$ is close to $\pi^{2}$. Linear growth rate: $\sigma=\mu-\pi^{2}$ is small.

$$
\underbrace{u_{t}}_{\epsilon \sigma}=u_{x x}+\mu(u-\frac{1}{6} \underbrace{u^{3}}_{\epsilon^{3}}+\cdots), \quad u=O(\epsilon)
$$

For a dominant balance between linear growth and nonlinearity, we expect

$$
\begin{aligned}
\epsilon \sigma & =\epsilon^{3} \\
\sigma & =O\left(\epsilon^{2}\right)
\end{aligned}
$$

This suggests the following expansion:

$$
\begin{aligned}
u & =\epsilon u_{1}(x, T)+\epsilon^{3} u_{3}(x, T)+O\left(\epsilon^{5}\right) \\
\mu & =\pi^{2}+\epsilon^{2} \mu_{2}+O\left(\epsilon^{4}\right) \\
T & =\epsilon^{2} t
\end{aligned}
$$

### 27.1 Model PDE Bifurcation Problem

$$
\begin{aligned}
u_{t} & =u_{x x}+\mu \sin u, \quad 0<x<1, t>0 \\
u(0, t) & =u(1, t)=0 \\
u(x, 0) & =f(x)
\end{aligned}
$$

- $u(x, t)=$ temperature
- $\mu=$ strength of the source
$u=0$ is
- linearly stable for $\mu<\pi^{2}$
- linearly unstable for $\mu>\pi^{2}$

Look at the effect of nonlinearity near the point of marginal stability, $\mu=\pi^{2}$. The dominant blance suggested

$$
\begin{aligned}
\mu-\pi^{2} & =O\left(\epsilon^{2}\right) \\
u & =O(\epsilon) \\
\text { time-scales } \quad t & =O\left(\frac{1}{\epsilon^{2}}\right)
\end{aligned}
$$

Expand:

$$
\begin{aligned}
\mu & =\pi^{2}+\epsilon^{2} \mu_{2}+O\left(\epsilon^{4}\right) \\
u & =\epsilon u_{1}(x, T)+\epsilon^{3} u_{2}(x, T)+O\left(\epsilon^{5}\right) \\
T & =\epsilon^{2} t \\
\partial_{t} & =\epsilon^{2} \partial_{T} \\
\epsilon^{2} u_{T} & =u_{x x}+\left(\pi^{2}+\epsilon^{2} \mu_{2}\right) \sin u, \quad 0<x<1, T>0 \\
u(0, t) & =u(1, t)=0 \\
\sin u & =u-\frac{1}{6} u^{3}+O\left(u^{5}\right) \\
& =\epsilon u_{1}+\epsilon^{3} u_{3}-\frac{1}{6} \epsilon^{3} u_{1}^{3}+O\left(\epsilon^{5}\right) \\
\epsilon^{3} u_{1, T}+\cdots & =\epsilon u_{1, x x}+\epsilon^{3} u_{3, x x}+\left(\pi^{2}+\epsilon^{2} u_{2}\right)\left(\epsilon u_{1}+\epsilon^{3}\left[u_{3}-\frac{1}{6} u_{1}^{3}\right]+\cdots\right) \\
u_{1, x x}+\pi^{2} u_{1} & =0 \\
u_{1}(0, t) & =u_{1}(1, t)=0 \\
O\left(\epsilon^{3}\right): \quad u_{3, x x}+\pi^{2} u_{3} & =u_{1, T}+\frac{\pi^{2}}{6} u_{1}^{3}-\mu_{2} u_{1} \\
u_{3}(0, t) & =u_{3}(1, t)=0
\end{aligned}
$$

We get

$$
\begin{aligned}
u_{1} & =a(T) \sin (\pi x) \\
u_{3, x x}+\pi^{2} u_{3} & =a_{1, T} \sin (\pi x)+\frac{\pi^{2}}{6} a^{3} \sin ^{3}(\pi x)-\mu_{2} a \sin (\pi x) \\
u_{3}(0, t) & =u_{3}(1, t)=0 \\
L u_{3} & =f(x) \\
L & =\frac{d^{2}}{d x^{2}}+\pi^{2}
\end{aligned}
$$

This is solvable if for $\phi$ such that $L \phi=0$, we have that

$$
\begin{aligned}
\left\langle\phi, L u_{3}\right\rangle & =\langle\phi, f\rangle \\
\left\langle L \phi, u_{3}\right\rangle & =\langle\phi, f\rangle \\
0 & =\langle\phi, f\rangle
\end{aligned}
$$

Thus, we must have that

$$
\begin{aligned}
a_{T} \underbrace{\left[\int_{0}^{1} \sin ^{2}(\pi x) d x\right]}_{=\frac{1}{2}}+\frac{\pi^{2}}{6} a^{3} \underbrace{\left[\int_{0}^{1} \sin ^{4}(\pi x) d x\right]}_{=\frac{3}{8}}-\mu_{2} a \underbrace{\left[\int_{0}^{1} \sin ^{2}(\pi x) d x\right]}_{=\frac{1}{2}} & =0 \\
\frac{1}{2} a_{T}+\frac{\pi^{2}}{16} a^{3}-\frac{1}{2} \mu_{2} a & =0 \\
a_{T}-\mu_{2} a+\frac{\pi^{2}}{8} a^{3} & =0
\end{aligned}
$$

This is typically called an amplitude equation (Laundau-Stuart). The equilibria are:

$$
a=0 \quad \text { OR } \quad a^{2}=\frac{8 \mu_{2}}{\pi^{2}}
$$



Figure 12: This is a (supercritical) pitchfork bifurcation. A rigorous analysis of the equilibrium states is obtained using Liapunov-Schmidt reduction.

Initial layer: take

$$
\begin{aligned}
t & =O(1) \\
\mu & =\pi^{2}+\epsilon^{2} \mu_{2} \\
u & =\epsilon u_{1}(x, t)+\epsilon^{3} u_{3}(x, t)+\cdots \\
u_{1, t} & =u_{1, x x}+\pi^{2} u_{1} \\
u_{1}(0, t) & =u_{1}(1, t)=0 \\
u_{1}(x, 0) & =f(x) \\
u_{1}(x, t) & =\sum_{n=1}^{\infty} c_{n} e^{-\left(n^{2}-1\right) \pi^{2} t} \sin (n \pi x) \\
& =c_{1} \sin (\pi x)+\sum_{n=2}^{\infty} c_{n} e^{-\left(n^{2}-1\right) \pi^{2} t} \sin (n \pi x) \\
c_{n} & =2 \int_{0}^{1} f(x) \sin (n \pi x) d x
\end{aligned}
$$

As $t \rightarrow \infty$,

$$
u_{1} \sim c_{1} \sin (\pi x)
$$

So we require

$$
\begin{aligned}
& a(T) \rightarrow c_{1} \quad \text { as } T \rightarrow 0 \\
& a(0)=2 \int_{0}^{1} f(x) \sin (\pi x) d x
\end{aligned}
$$

Final: Tuesday June 12 from 1:30-3:30
Office Hours: Monday 2:30-4:00

### 28.1 Outline of Topics

1. Dimensional analysis and scaling

- Buckingham-Pi Theorem
- Self-similarity

2. Asymptotic expansions

- o, $O$ notation
- Asymptotic vs. convergent series
- Expansion of integrals
- (Did NOT cover the method of stationary phase or steepest descent)

3. Regular vs. singular perturbation problems

- Algebraic equations (e.g. polynomials)
- Dominant balance (distinguished limits)

4. Method of matched asymptotics

- Construct inner \& outer solutions and match them
- Uniform solutions
- Initial layer problems (e.g. enzyme dynamics)
- Slow-fast dynamics in systems of ODE's
- Boundary layer problems

5. Method of multiple scales

- Poincaré-Lindstedt method (periodic solutions)
- Multiple scales $(t, T)$ and applications to oscillations
- Method of averaging
- WKB method
- Fredholm alternative \& solvability conditions $\Rightarrow$ these were a unifying theme

The final will probably be 5 questions (roughly one from each topic).

1. Multiple scales
2. Boundary layers
3. Nondimensionalization
4. Asymptotics

For example:

- Nondimensionalize this equation
- Here's a polynomial involving $\epsilon$, find the roots

Most of this is discussed in chapters 1 and 2 of Applied Mathematics.
Things to know:

- Taylor expansion for tan


### 28.2 Sample Problems

Example 28.1. Logan 2.1.4

$$
\begin{aligned}
f(y, \epsilon) & =\frac{1}{(1+\epsilon y)^{3 / 2}} \\
y & =y_{0}+\epsilon y_{1}+O\left(\epsilon^{2}\right)
\end{aligned}
$$

Expand $f(y, \epsilon)$ in $\epsilon$ up to $O\left(\epsilon^{2}\right)$.

$$
\begin{aligned}
f(y, \epsilon) & =(1+\epsilon y)^{-3 / 2} \\
& =1-\frac{3}{2} \epsilon y+\frac{1}{2}\left(-\frac{3}{2}\right)\left(-\frac{5}{2}\right)(\epsilon y)^{2}+O\left(\epsilon^{3}\right) \\
& =1-\frac{3}{2} \epsilon y+\frac{15}{8} \epsilon^{2} y^{2}+O\left(\epsilon^{3}\right) \\
& =1-\frac{3}{2} \epsilon y_{0}+\epsilon^{2}\left[\frac{15}{8} y_{0}^{2}-\frac{3}{2} y_{1}\right]+O\left(\epsilon^{3}\right)
\end{aligned}
$$

Example 28.2. Logan 2.1.5h

How does $\exp (\tan \epsilon)$ behave as $\epsilon \rightarrow 0$ ? We are supposed to show that $\exp (\tan \epsilon)=O(1)$.

$$
\begin{array}{rll}
f(\epsilon)=O(g(\epsilon)) & \Rightarrow & |f(\epsilon)| \leq C|g(\epsilon)| \quad \text { for }|\epsilon|<\delta \\
f(\epsilon)=o(g(\epsilon)) & \Rightarrow & \left|\frac{f(\epsilon)}{g(\epsilon)}\right| \rightarrow 0 \quad \text { as } \epsilon \rightarrow 0 \quad(\text { if } g(\epsilon) \neq 0) \\
f(\epsilon) \sim g(\epsilon) & \Rightarrow & \left|\frac{f(e)}{g(\epsilon)}\right| \rightarrow 1
\end{array}
$$

$\sim$ and $o$ each imply $O$

$$
\begin{aligned}
f(\epsilon) & =\sin \left(\frac{1}{\epsilon}\right) \\
g(\epsilon) & =1 \\
f & =o(g) \quad \text { as } \epsilon \rightarrow 0 \quad(c=1)
\end{aligned}
$$

$$
\begin{aligned}
\exp (\tan \epsilon) \sim 1 & \text { as } \epsilon \rightarrow 0 \\
\exp (\tan \epsilon)-1 \sim \epsilon & \text { as } \epsilon \rightarrow 0
\end{aligned}
$$

$$
\begin{aligned}
\exp (\tan \epsilon) & =\exp \left(\epsilon+O\left(\epsilon^{3}\right)\right) \\
& =1+\left(\epsilon+O\left(\epsilon^{3}\right)\right)+O\left(\epsilon^{2}\right) \\
& =1+O(\epsilon) \\
\lim _{\epsilon \rightarrow 0} \exp (\tan \epsilon) & =1
\end{aligned}
$$

$$
\exists \delta>0 \quad \text { such that } \quad|\exp (\tan \epsilon)-1| \leq 1 \quad \text { for }|\epsilon|<\delta
$$

$$
|\exp (\tan \epsilon)| \leq 2 \cdot 1 \quad \text { for }|\epsilon|<\delta
$$

## Example 28.3. Logan 1.2.3

$$
m^{\prime}=a x^{2}-b x^{3}
$$

- $m=$ biomass
- $x=$ linear dimension
- $a x^{2}$ is the growth term (proportional to the surface area)
- $b x^{3}$ is the eating term (proportional to the volume)

Assume $m=\rho x^{3}$.

$$
\begin{aligned}
3 \rho x^{2} x^{\prime} & =a x^{2}-b x^{3} \\
x(0) & =x_{0}
\end{aligned}
$$

Nondimensionalize.
The dimensions are

- $M=$ biomass
- $L=$ length
- $T=$ time

The parameters are

- $a,[a]=\frac{M}{T L^{2}}$
- $b,[b]=\frac{M}{T L^{3}}$
- $\rho,[\rho]=\frac{M}{L^{3}}$
- $x_{0},\left[x_{0}\right]=L$

The variables are

- $t,[t]=T$
- $x,[x]=L$

We have 3 dimensions and 4 parameters, so we should have 1 dimensionless parameter. Let's leave $x_{0}$ alone and use $a, b, \rho$ to nondimensionalize mass, length, and time.

$$
\begin{aligned}
{\left[\frac{a}{b}\right] } & =L \\
{\left[\rho \frac{a^{3}}{b^{3}}\right] } & =M \\
{\left[\frac{\rho}{b}\right]=} & \\
x^{*} & =\frac{x}{a / b} \\
t^{*} & =\frac{t}{\rho / b} \\
\text { (think) } \quad 3\left(x^{*}\right)^{2}\left(x^{*}\right)^{\prime} & =\left(x^{*}\right)^{2}-\left(x^{*}\right)^{3} \\
x^{*}(0) & =\frac{b x_{0}}{a}
\end{aligned}
$$

## Index

amplitude equation, 79
Fredholm alternative, 60
regular, 17
secular terms, 56
singular, 17, 19
turning point, 44, 75

